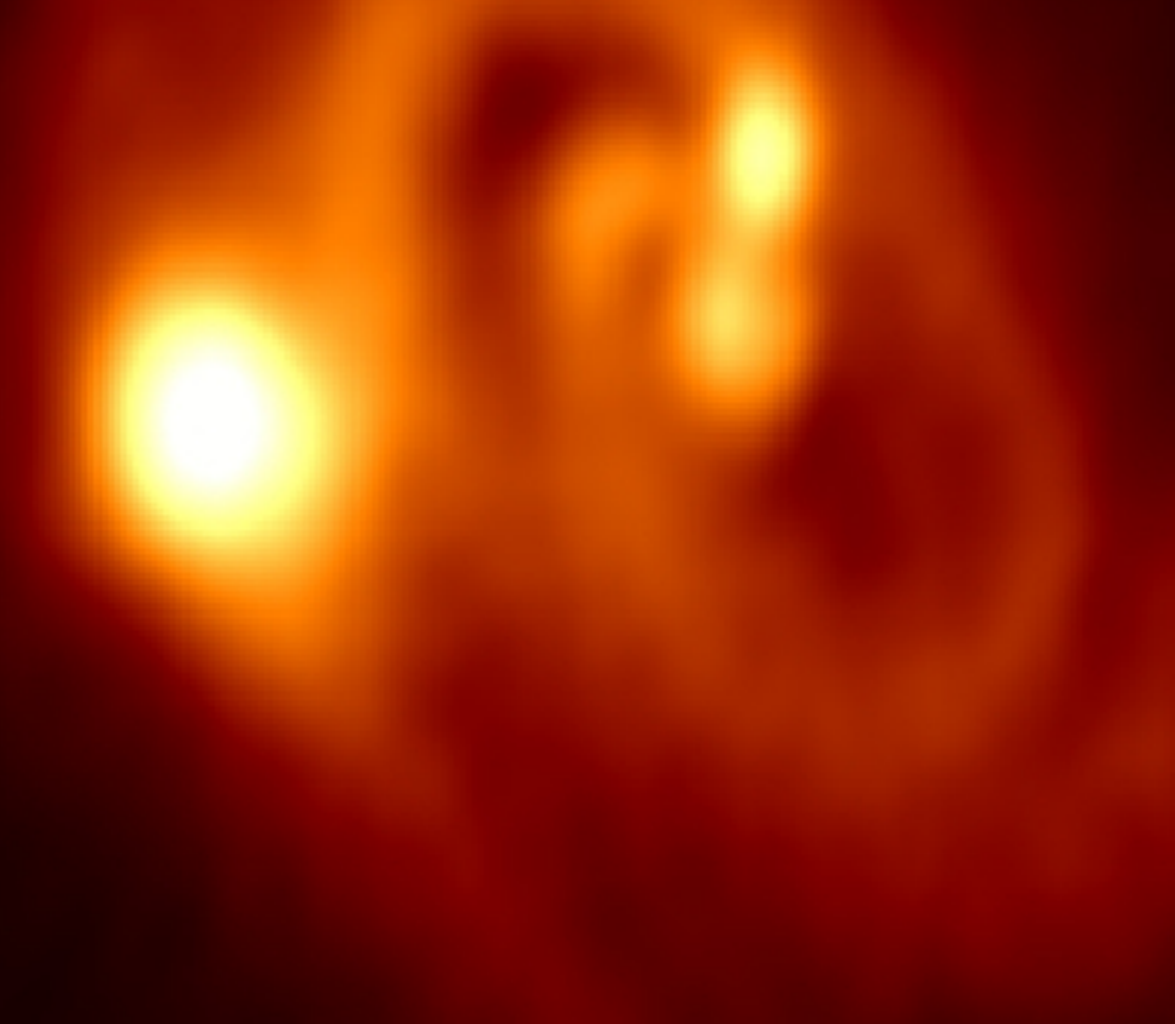


THE STRUCTURES OF EMBEDDED DISKS WITH ALMA



PATRICK SHEEHAN

CIERA FELLOW

GREAT BARRIERS IN PLANET FORMATION

JULY 22, 2019

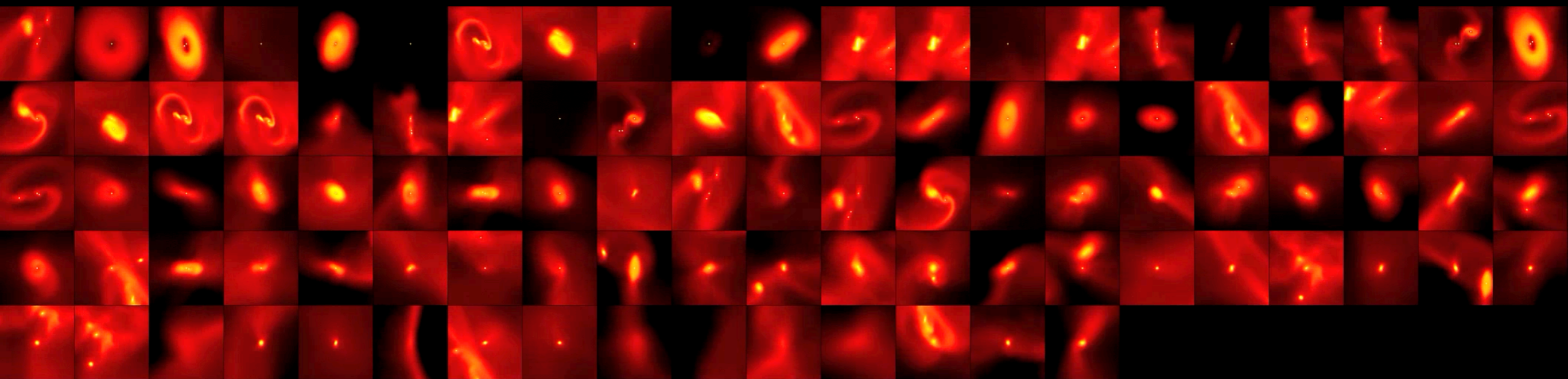
COLLABORATORS:

JOHN TOBIN

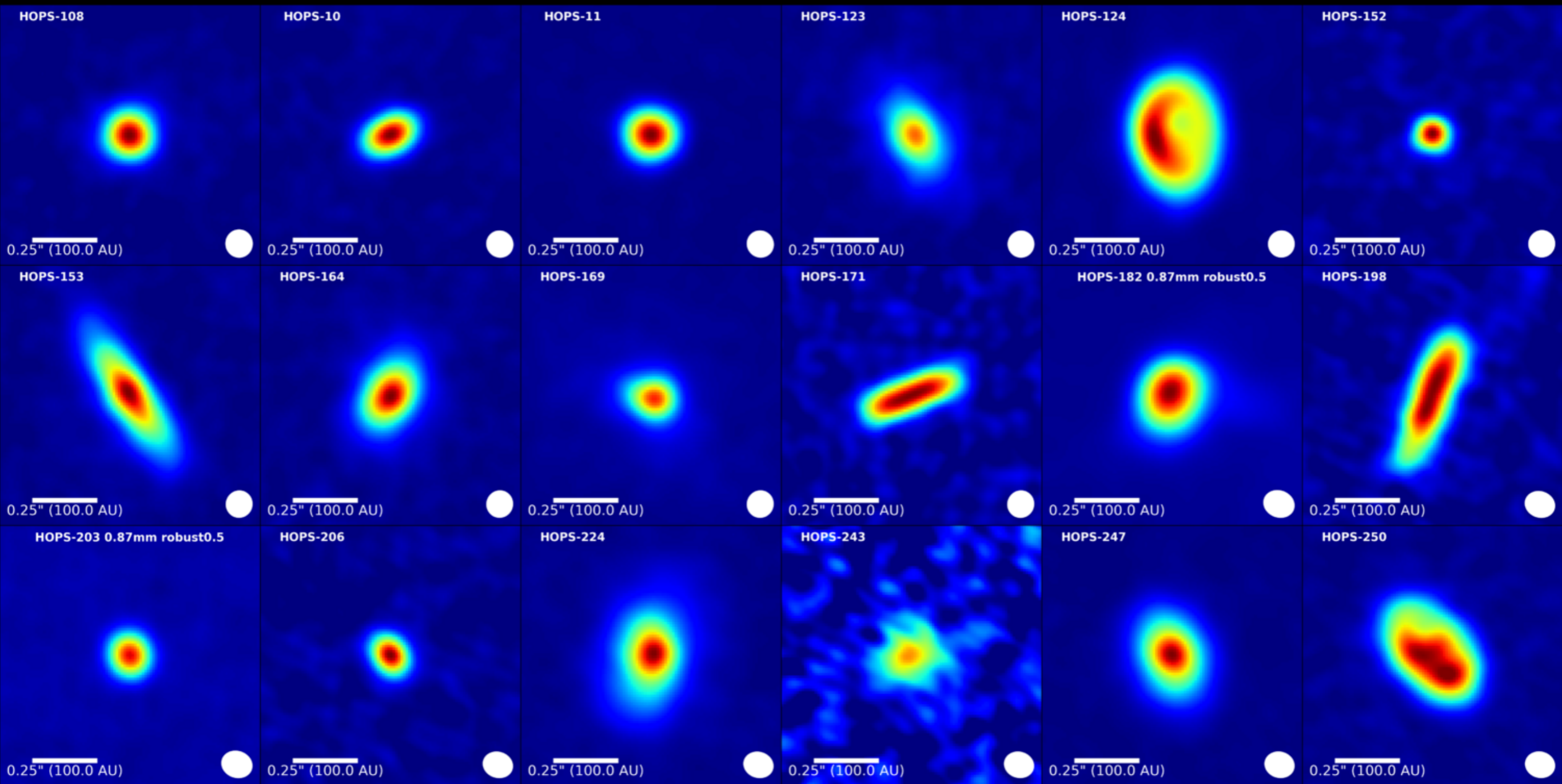
JOSH EISNER

Image credit: Reynolds et al., in prep.

Protostellar disk formation

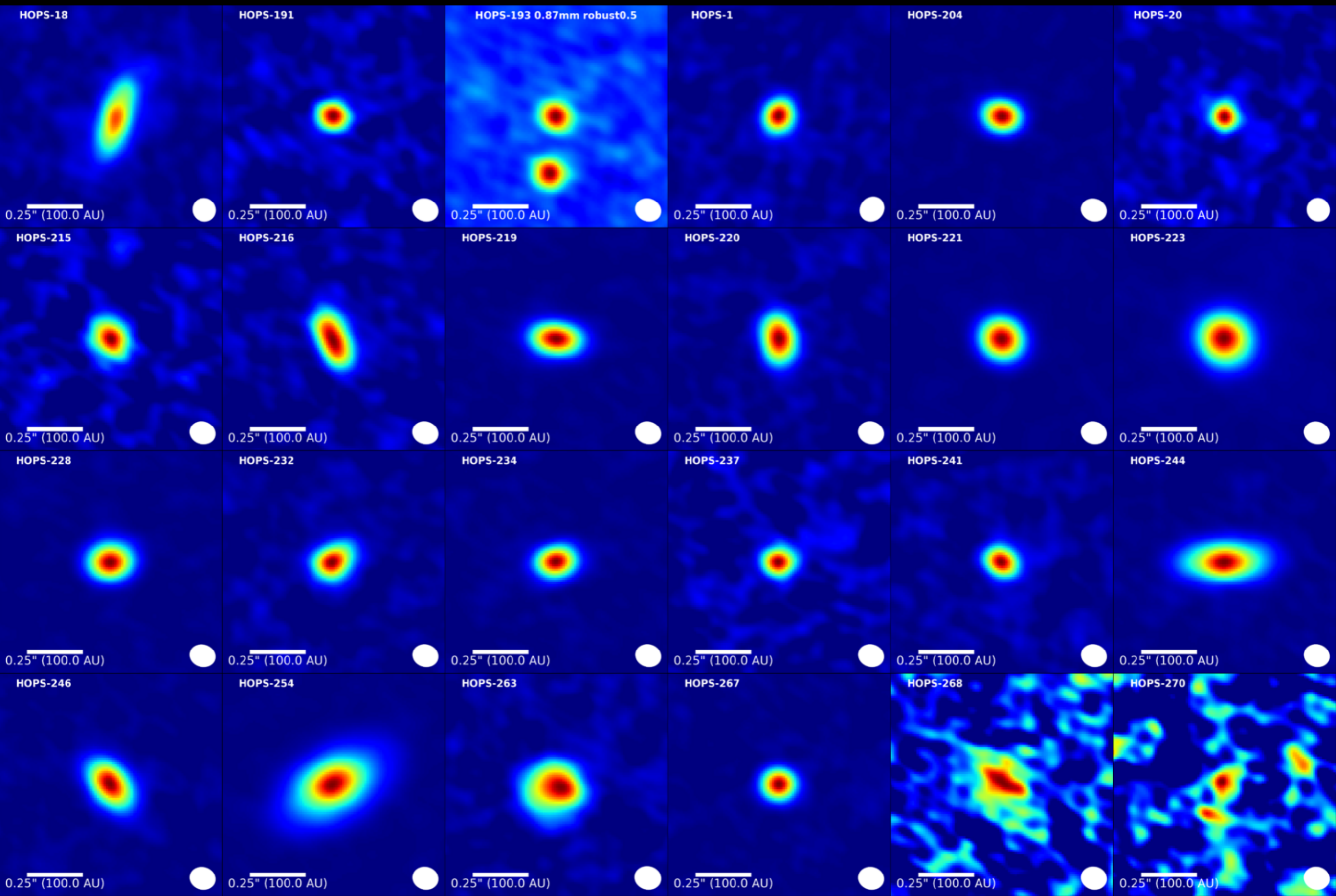


VANDAM: Orion Class 0 - ALMA

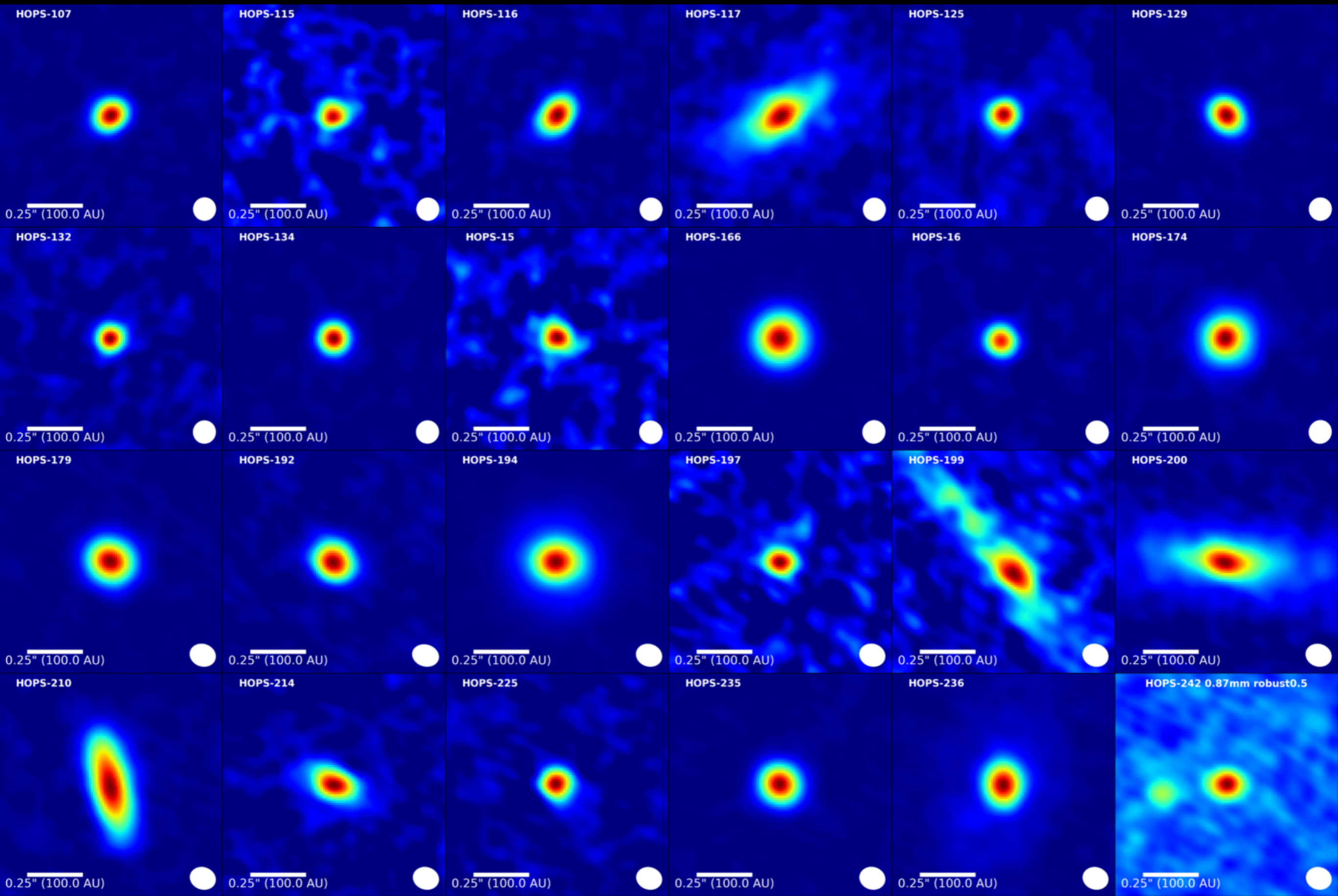


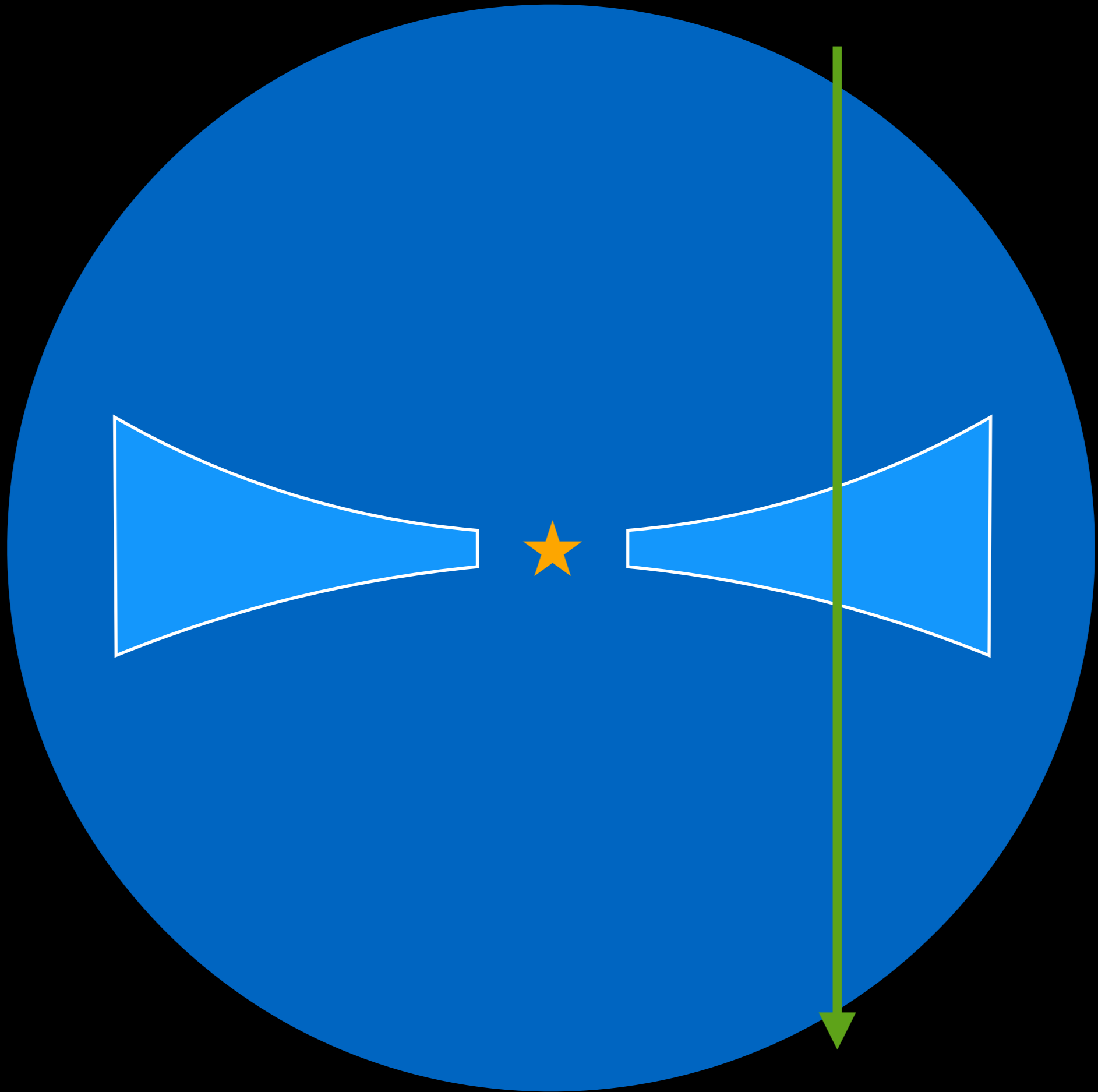
- >300 protostellar disks in Orion with ~40 au resolution at 345 GHz
- ~100 Class 0's, ~200 Class I's/Flat Spectrum
- PI: John Tobin

VANDAM: Orion Class I



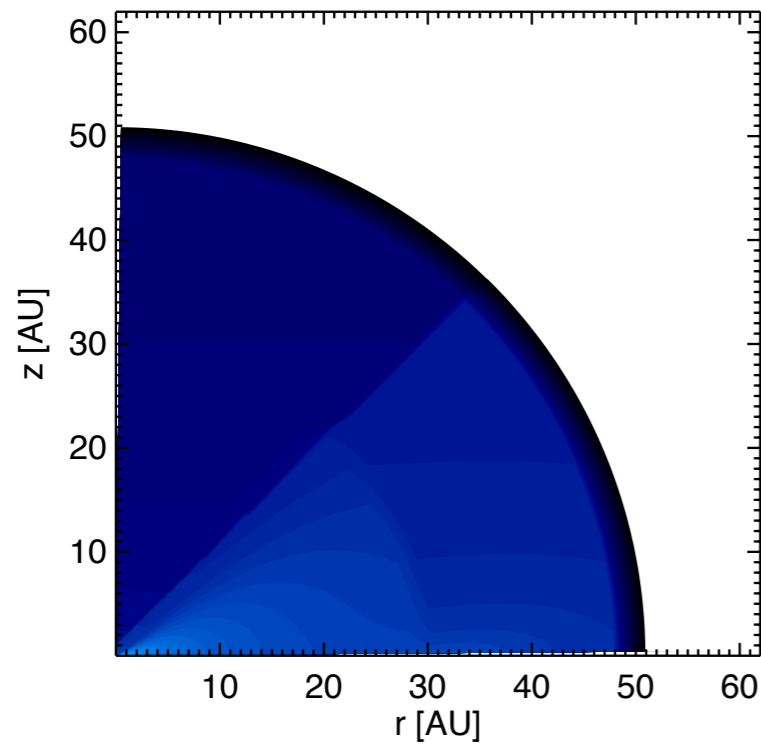
VANDAM: Orion Flat Spectrum/Class I





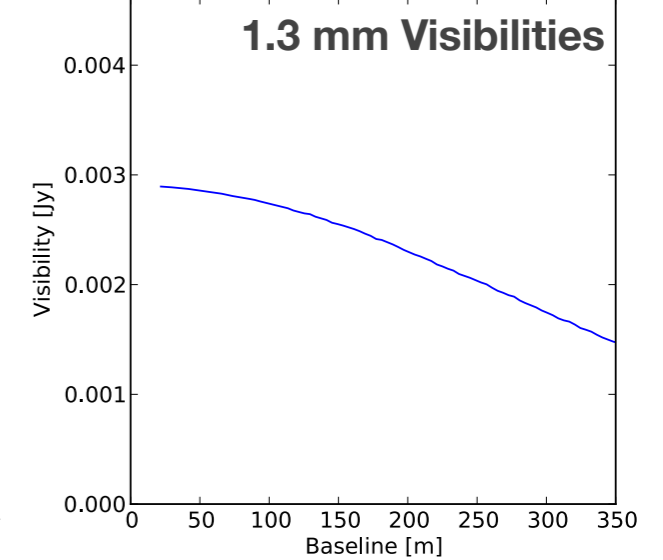
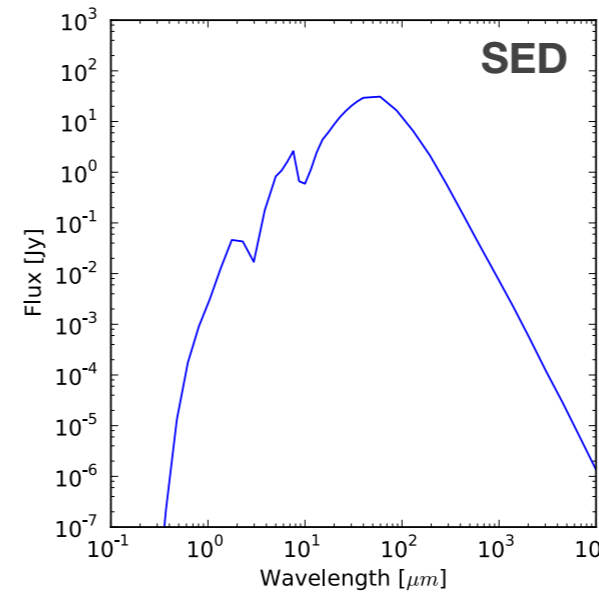
Radiative transfer forward modeling

parametric density distribution
disk + envelope + outflow cavity



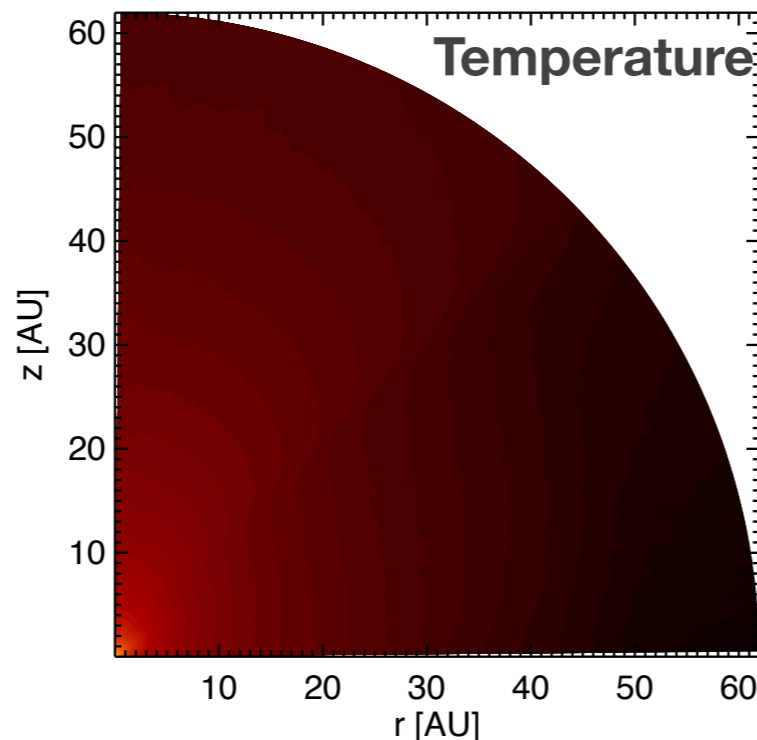
RADMC-3D

synthetic observations

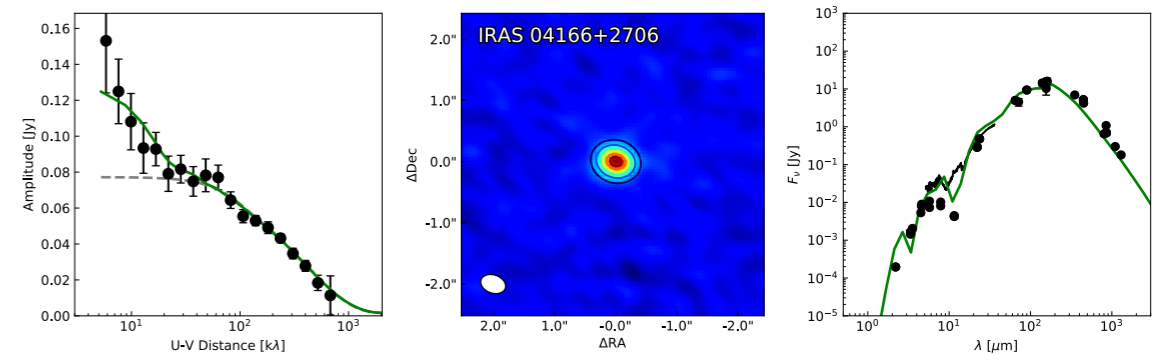


emcee

RADMC-3D



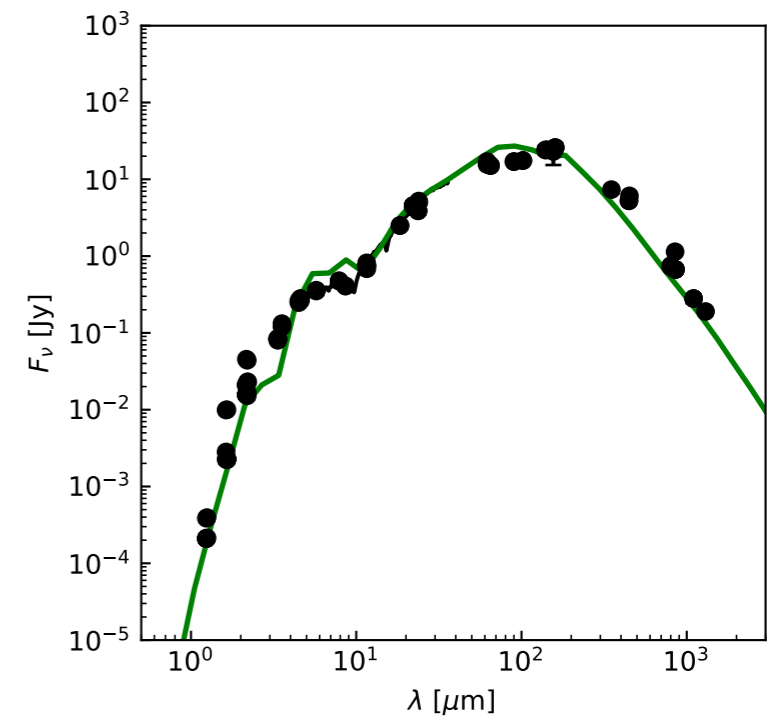
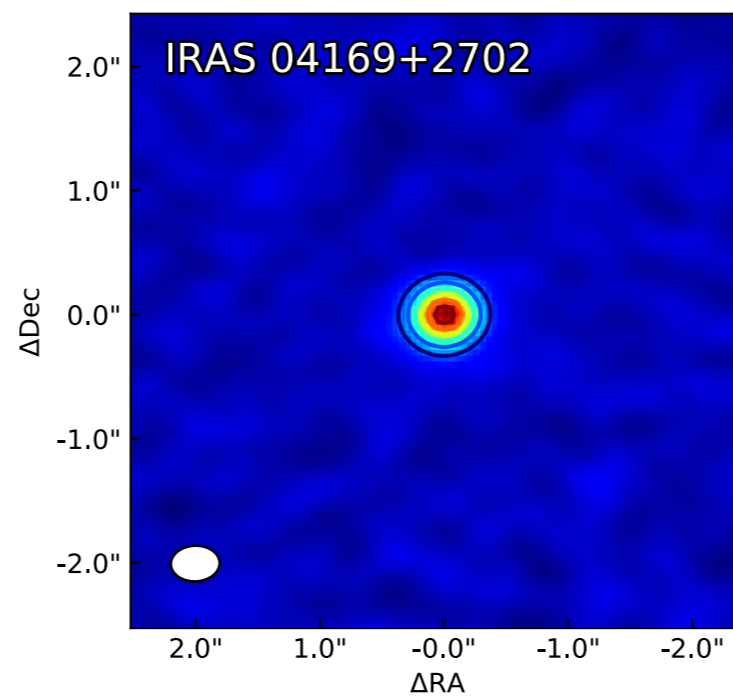
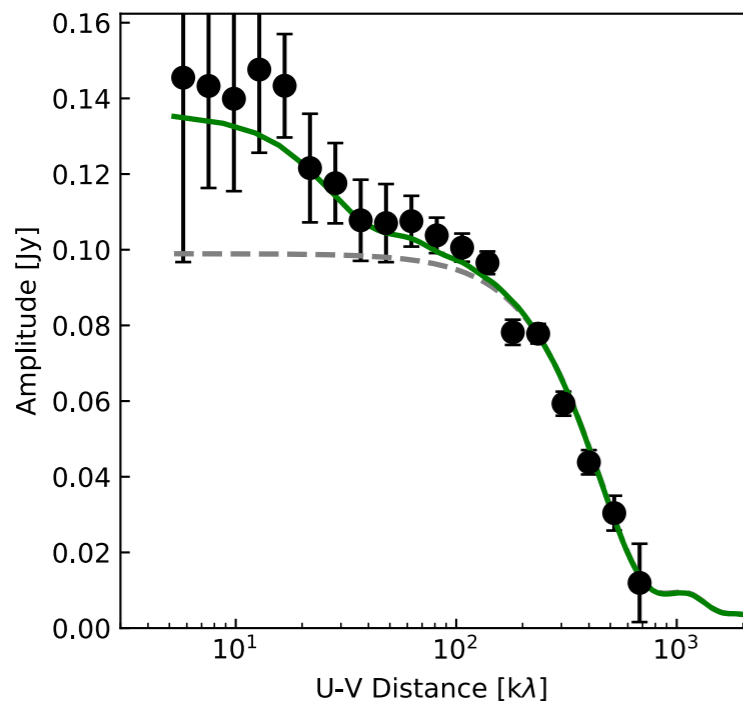
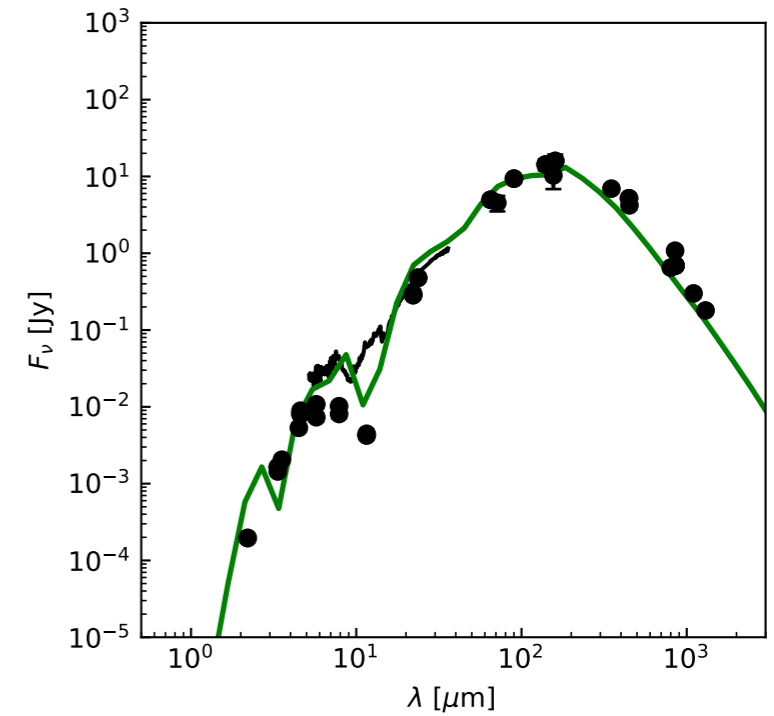
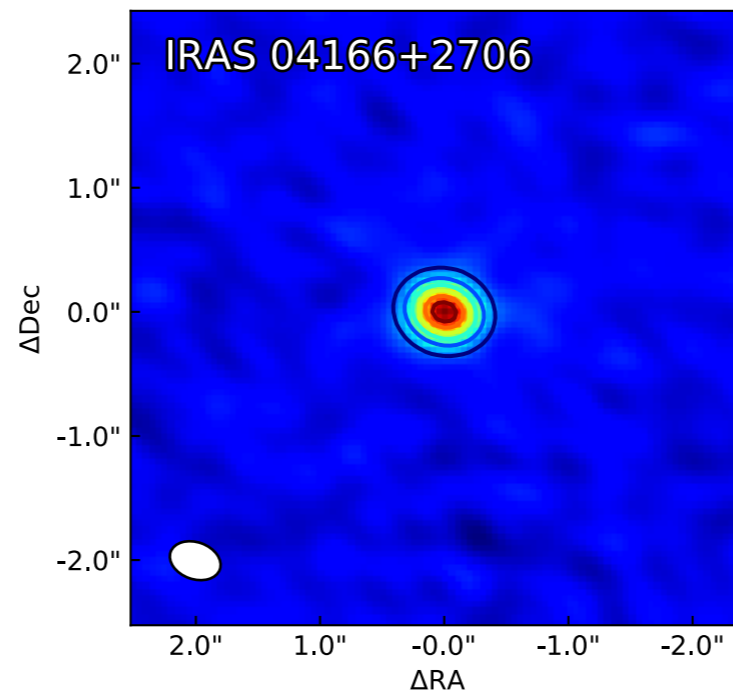
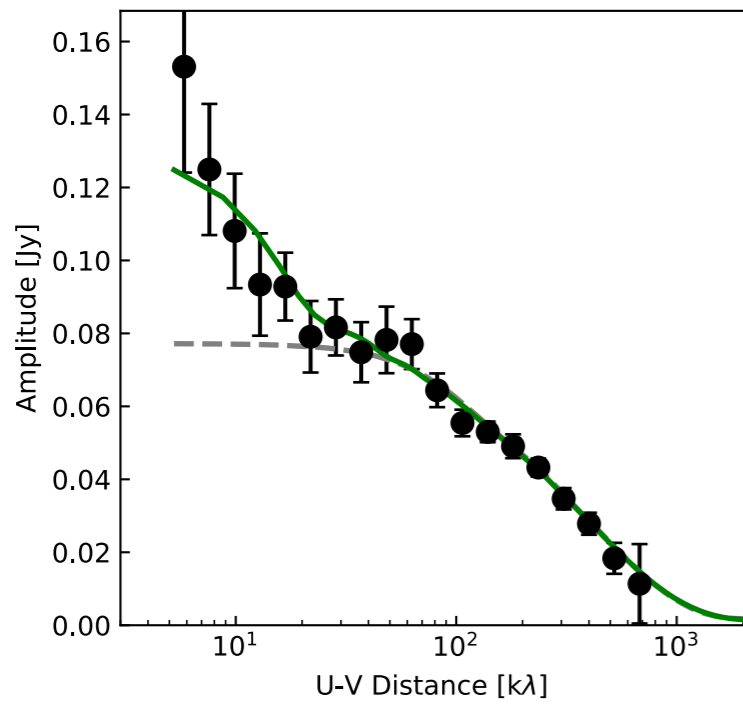
best fit model



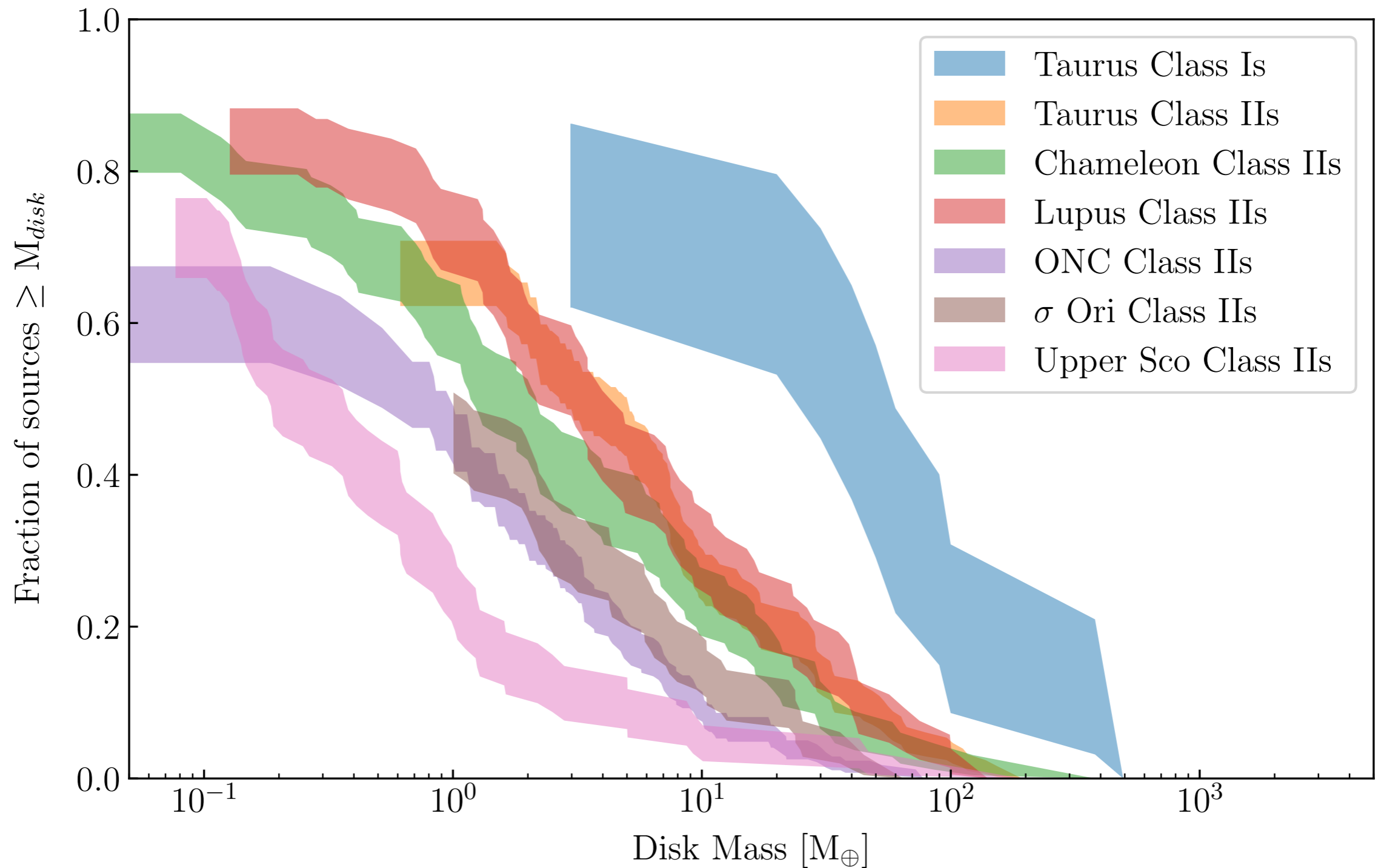
pdspy: Available on GitHub now! Documentation sparse, but more coming soon.

Class I protostars in Taurus

Sheehan & Eisner 2017b



The Initial Mass Budget for Forming Planets?

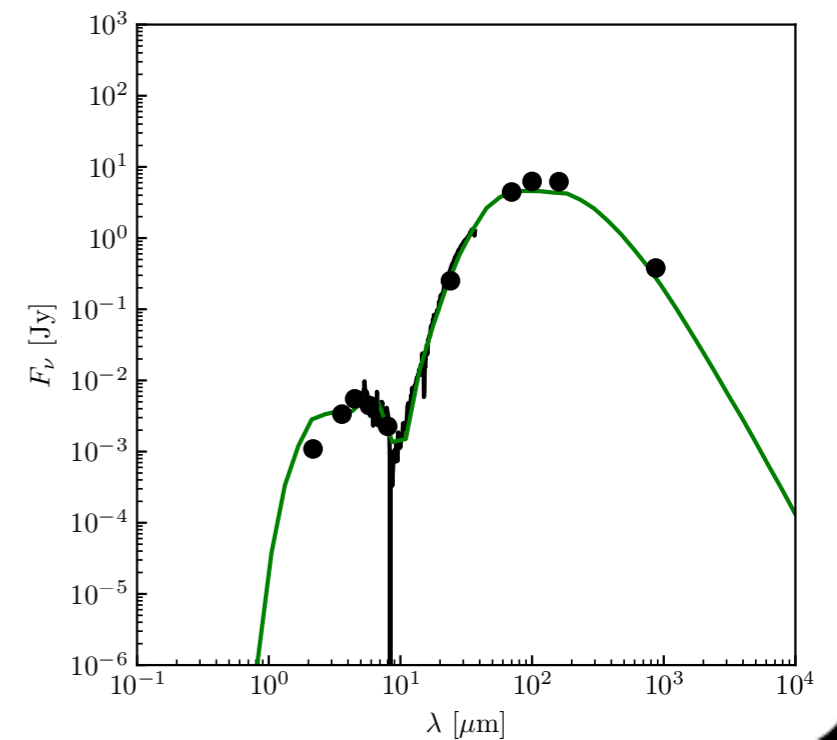
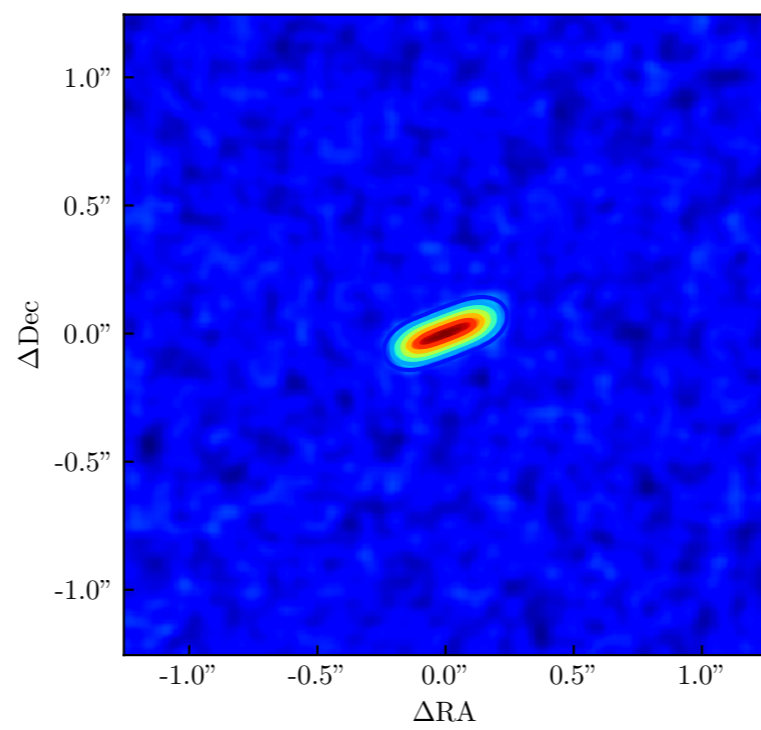
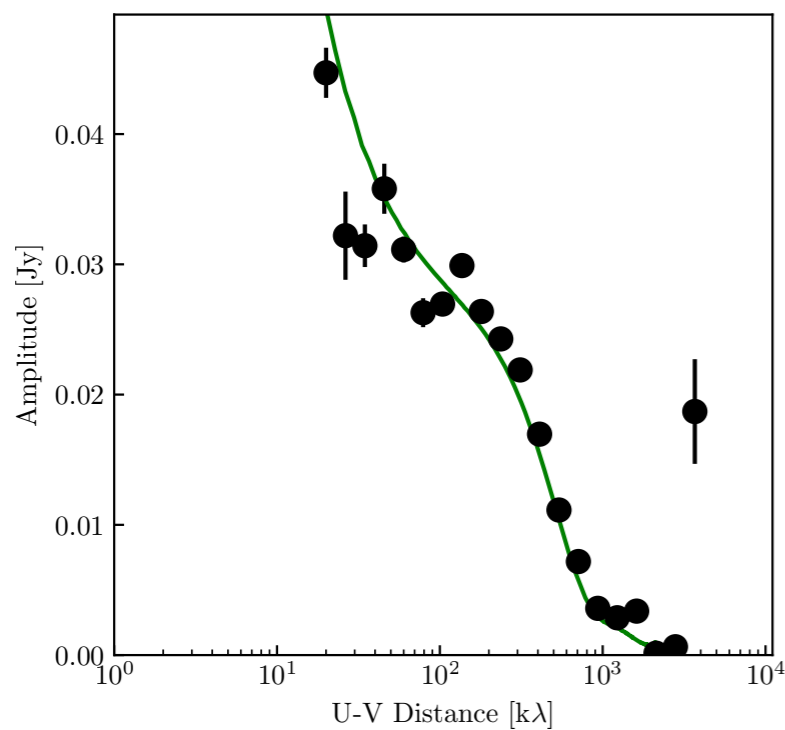
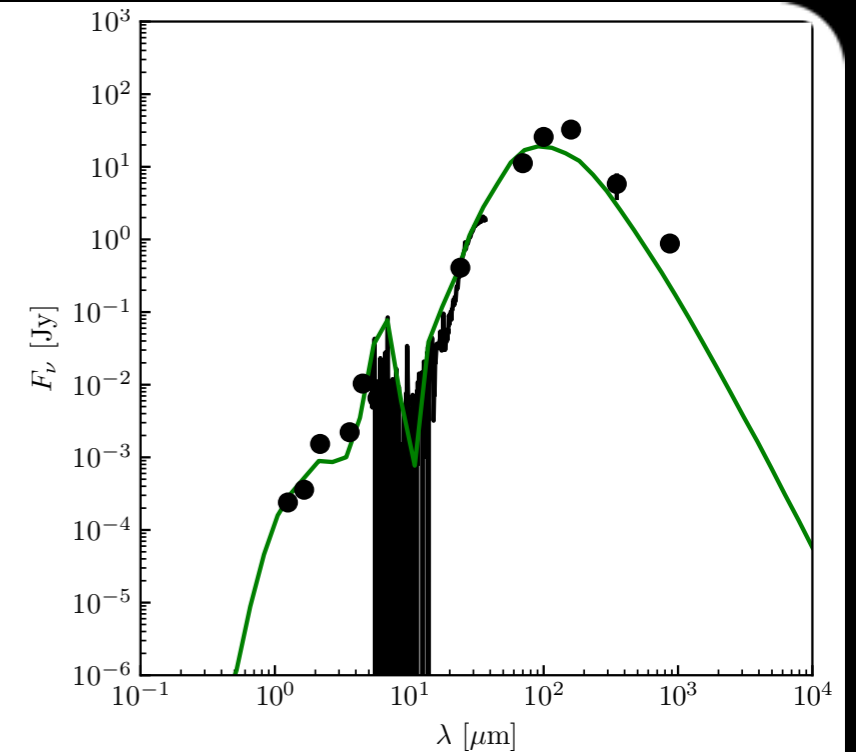
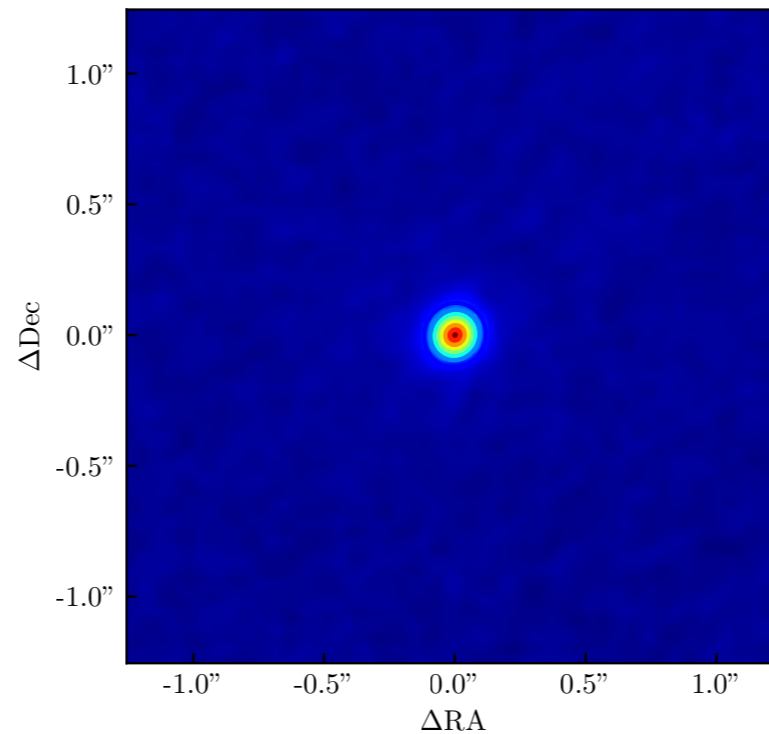
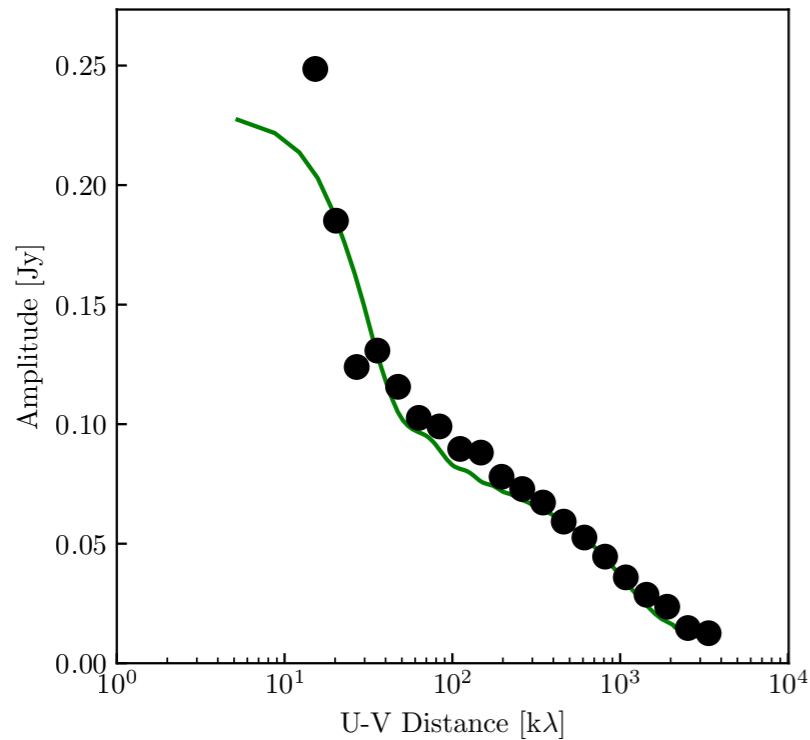


Sheehan & Eisner 2017b

- With 10 sources, only a $\sim 2.5\sigma$ result \Rightarrow need a bigger sample!

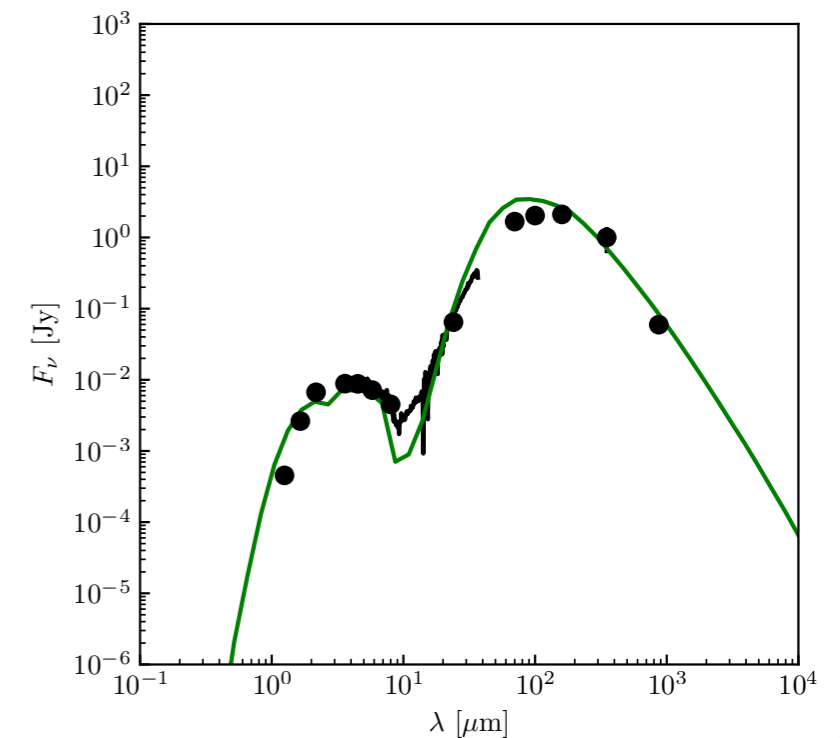
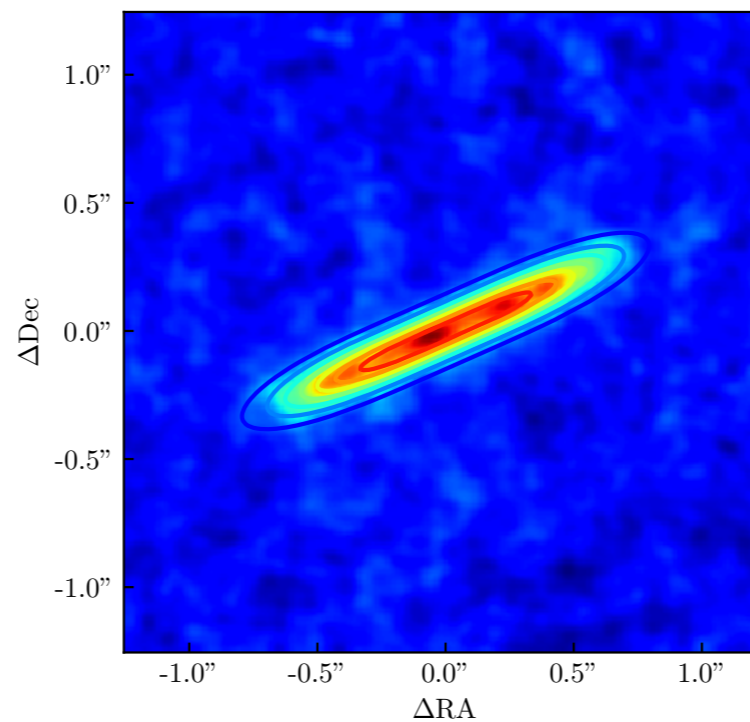
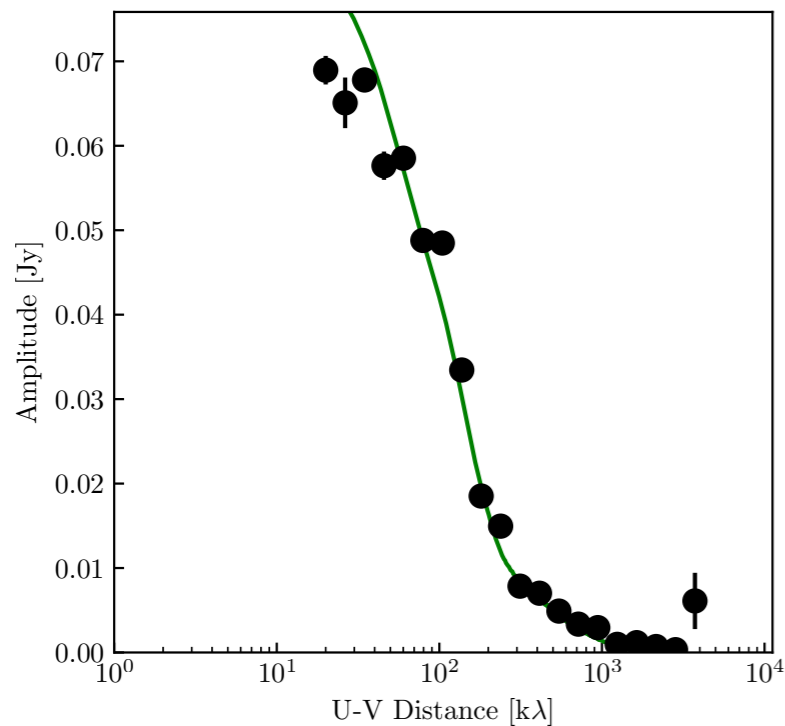
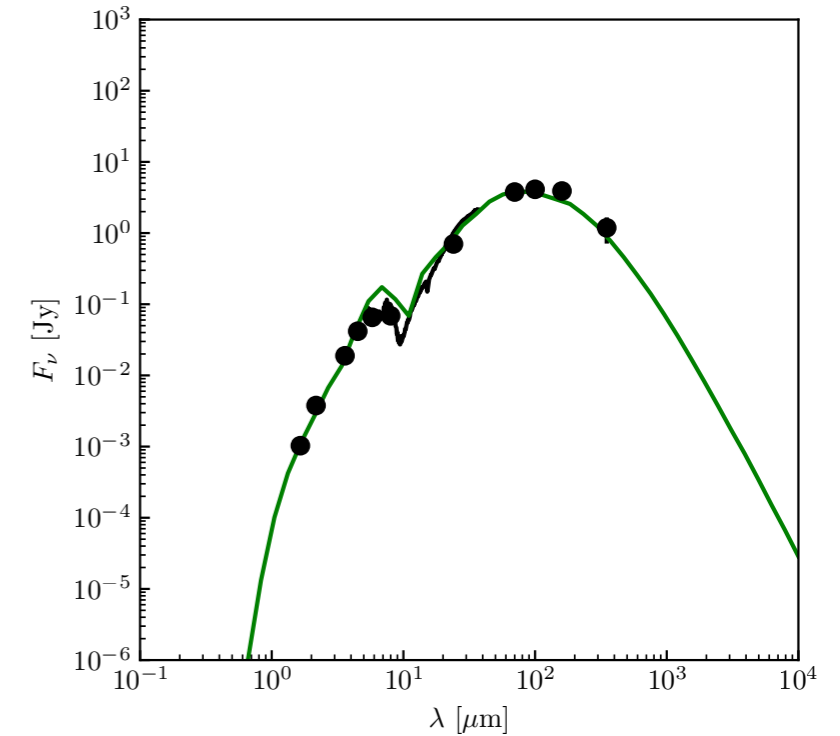
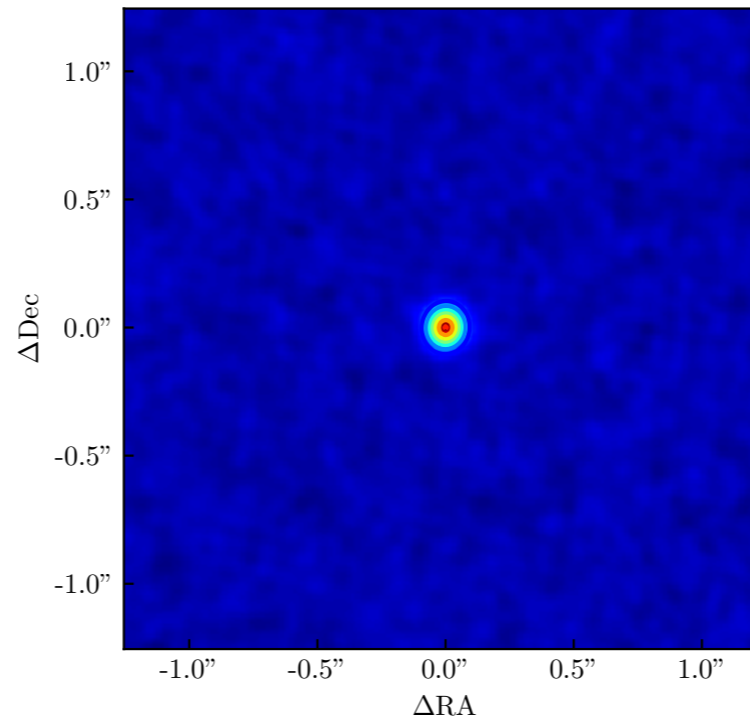
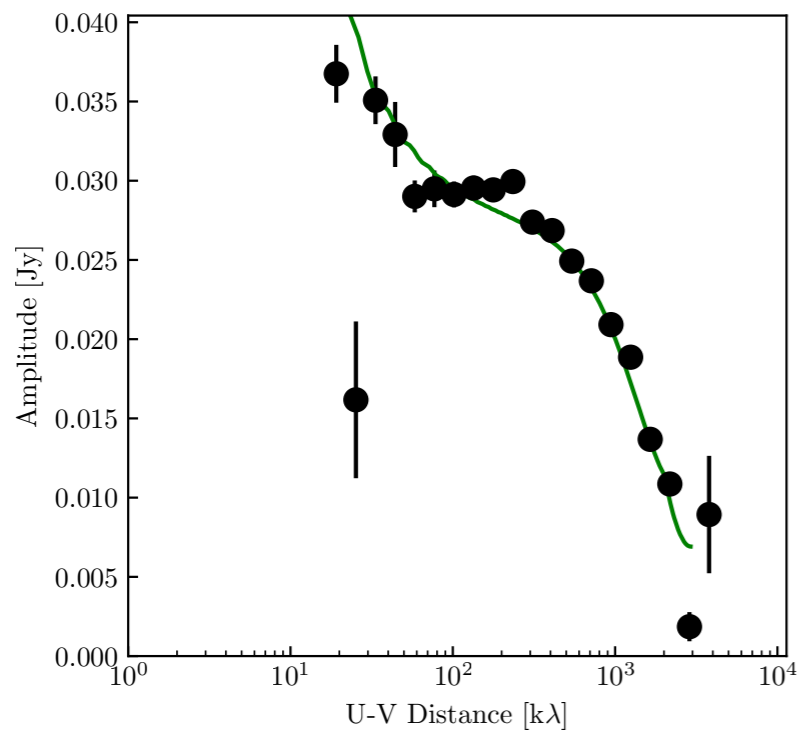
Modeling Class protostellar disks in Orion

- 5,000,000 core-hour allocation to model all 200 single protostars in the VANDAM: Orion survey with NSF XSEDE supercomputers



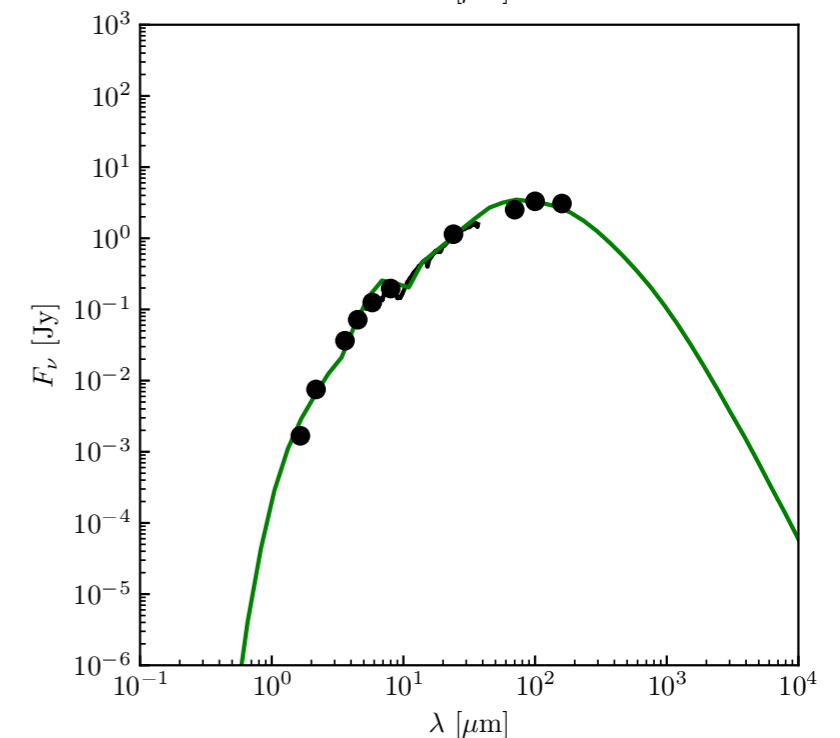
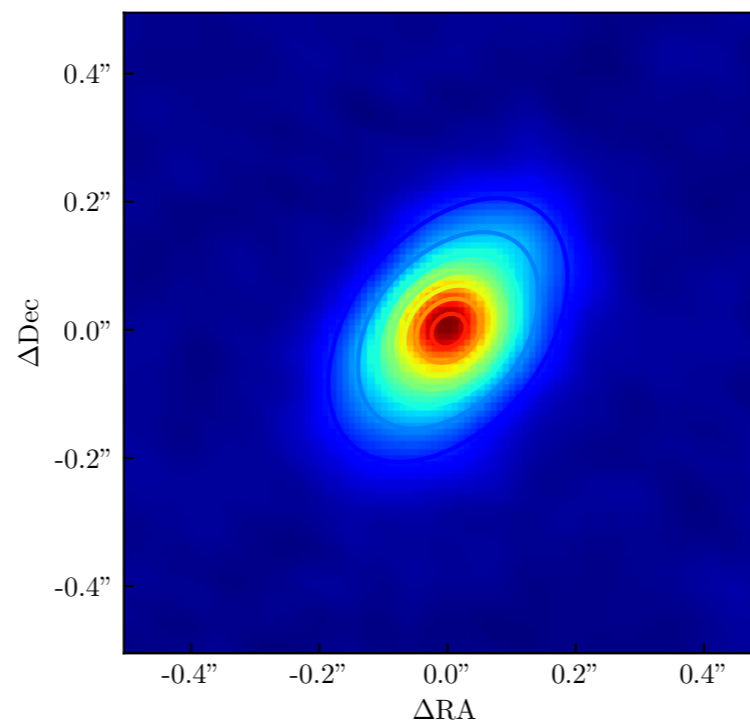
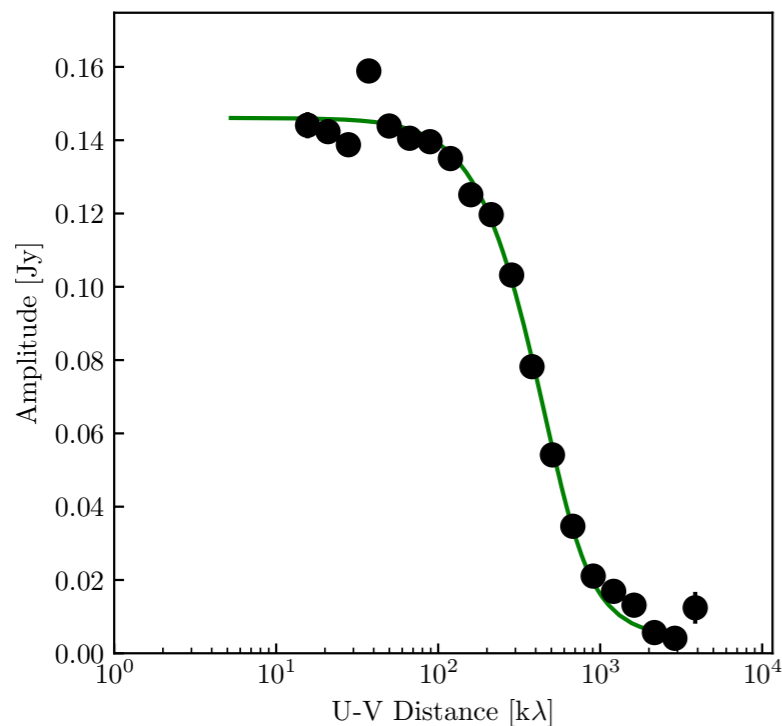
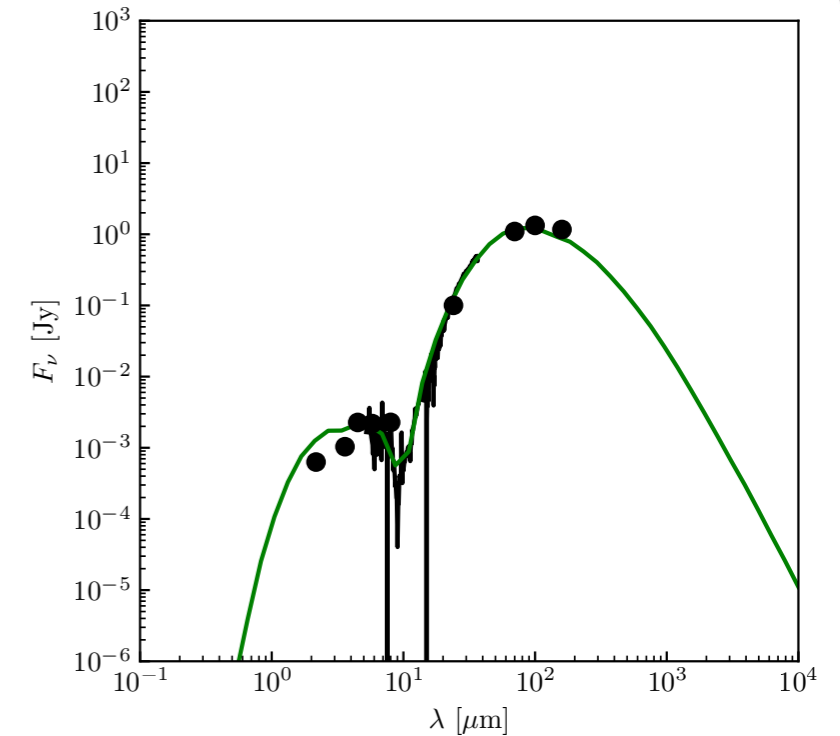
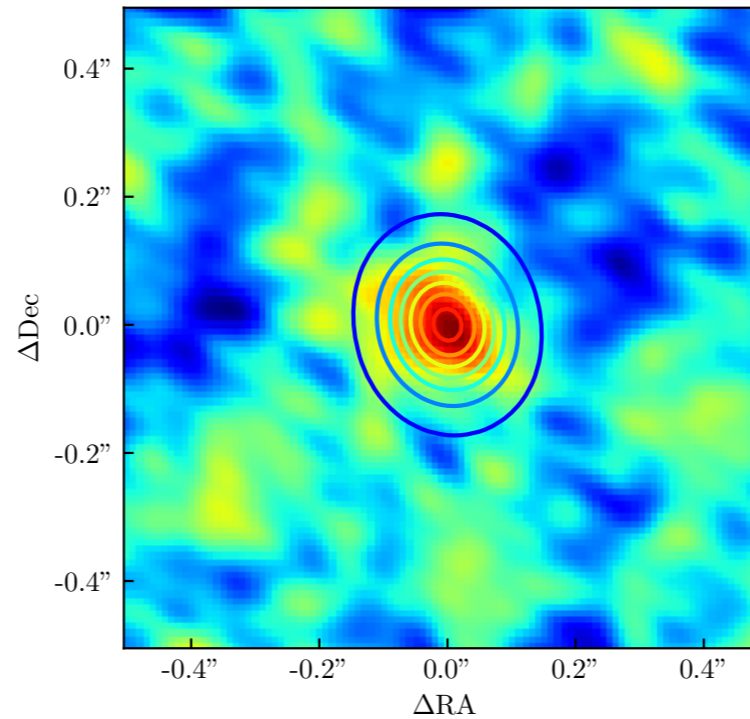
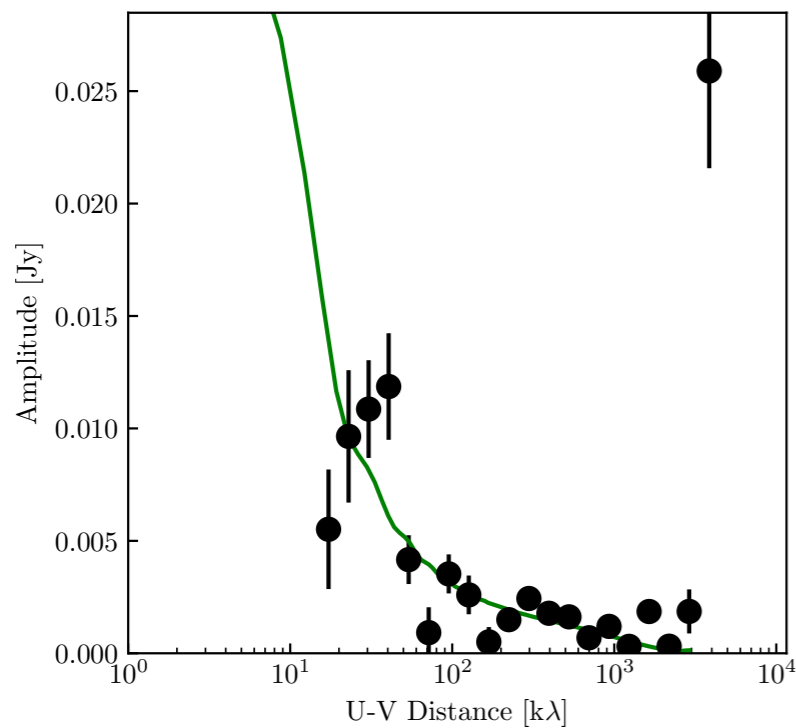
Modeling Class protostellar disks in Orion

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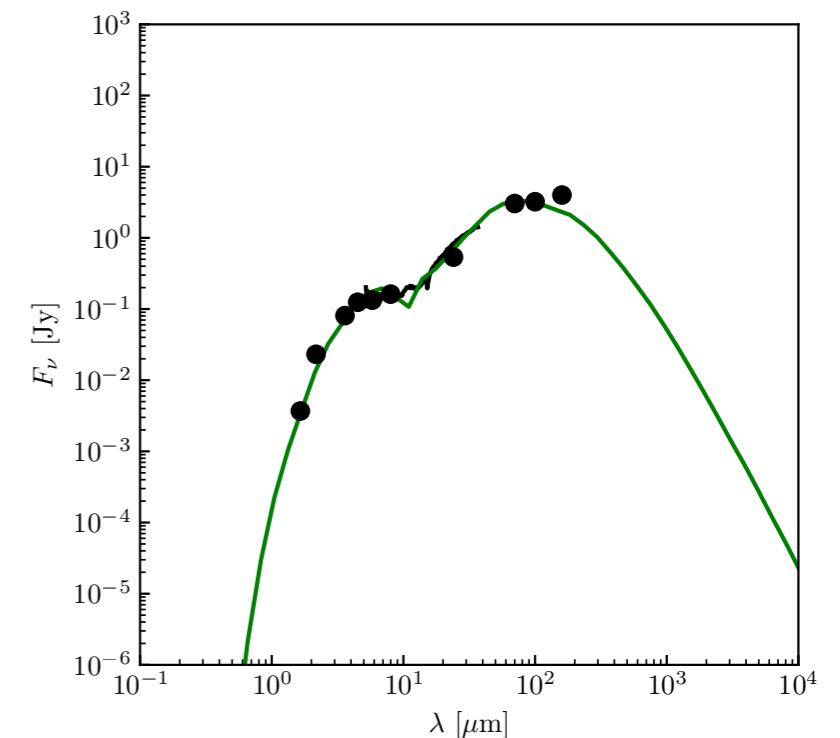
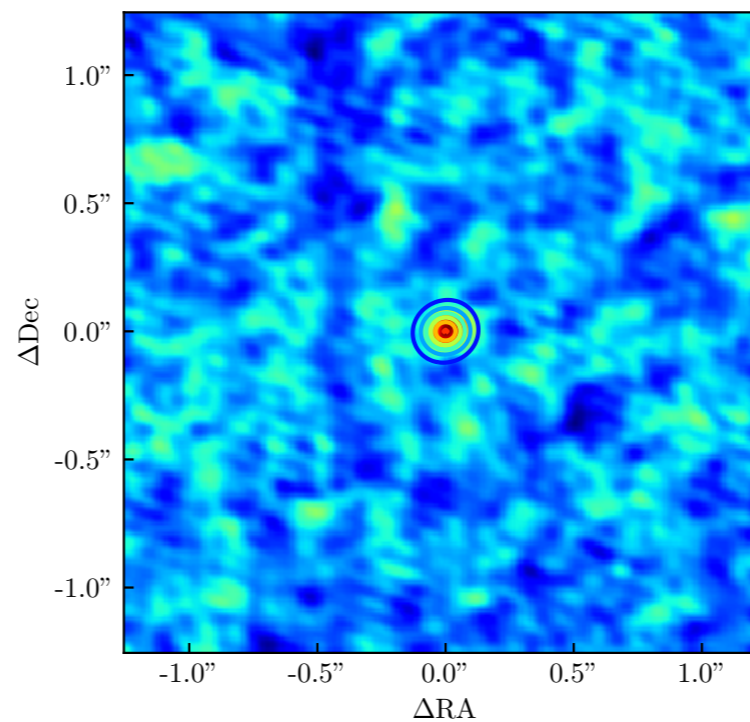
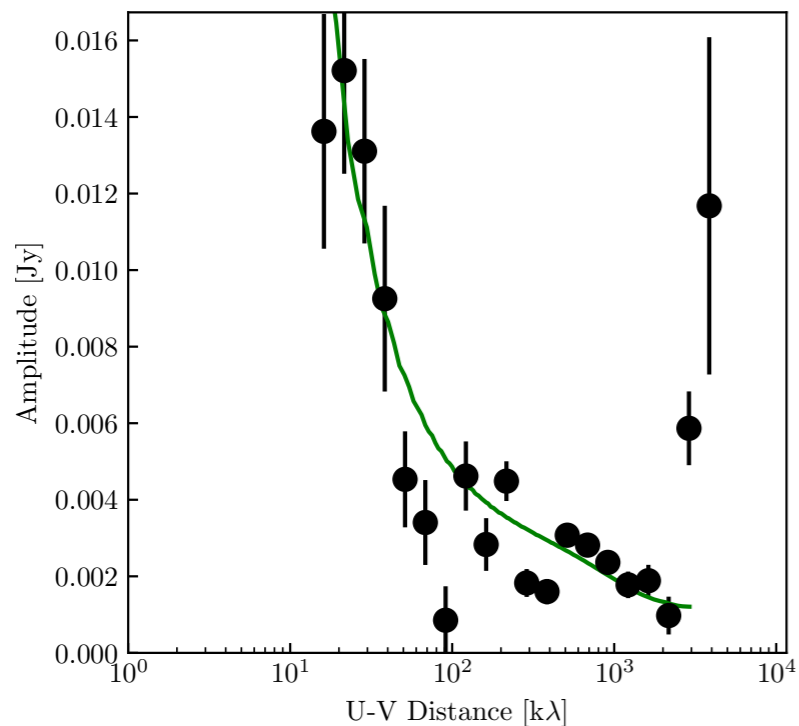
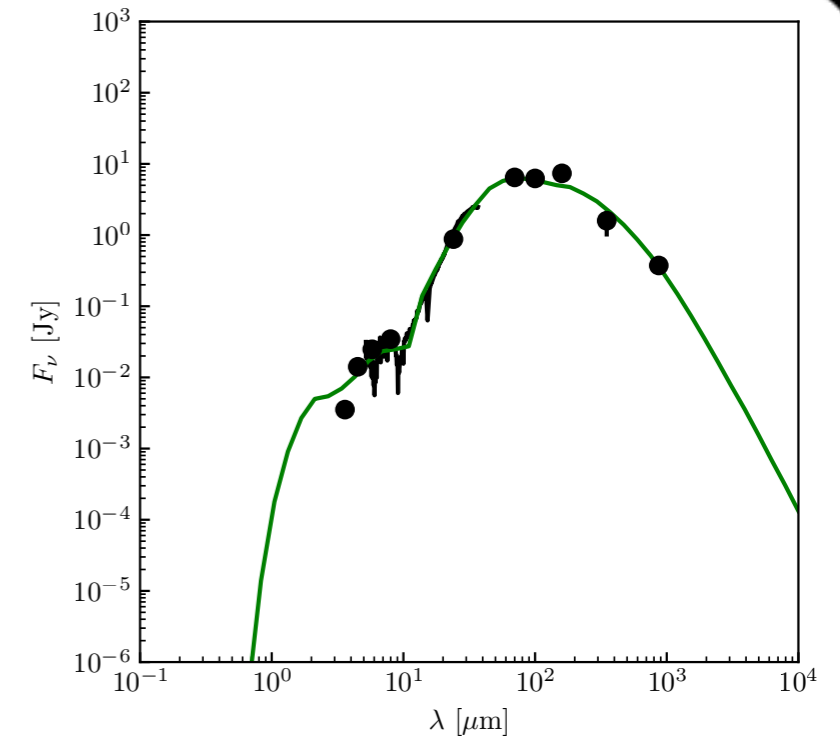
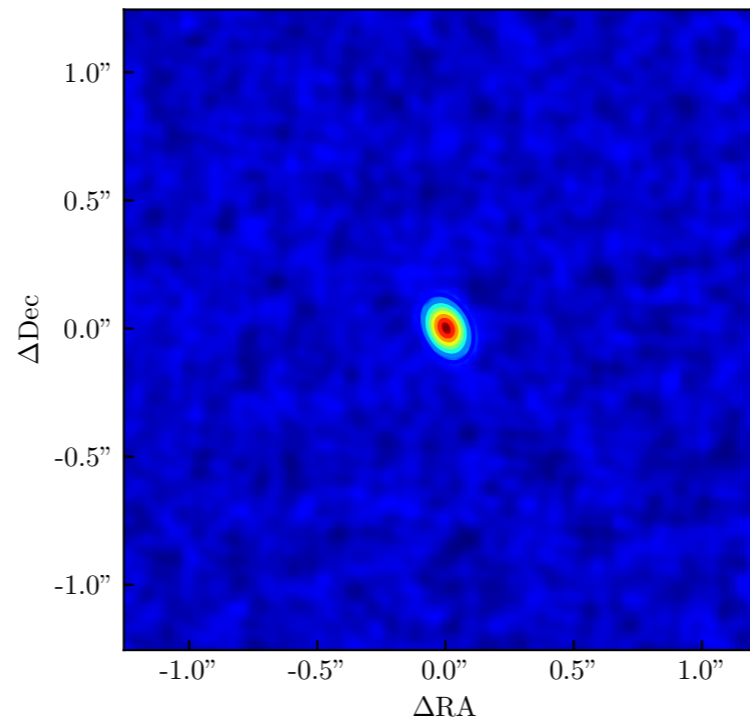
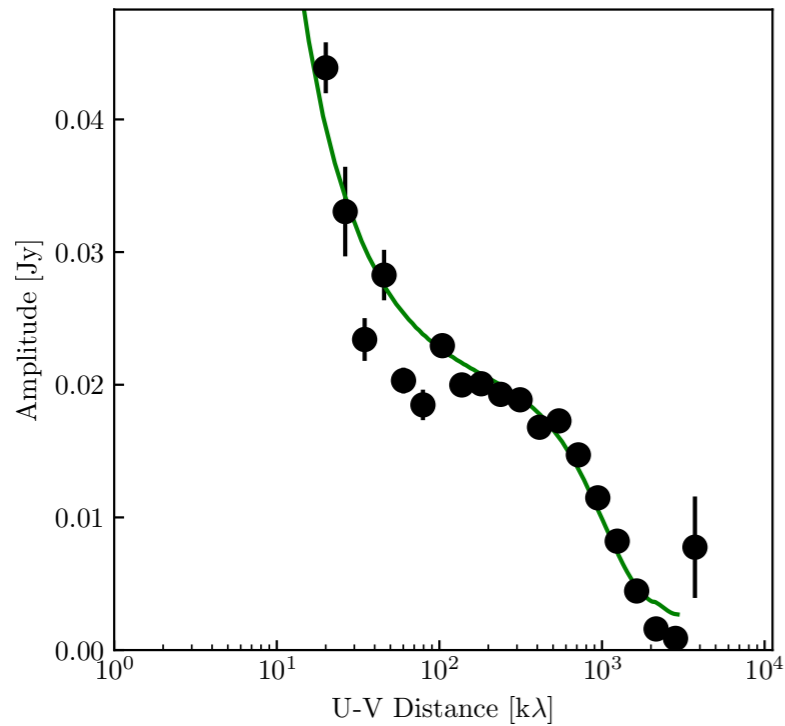
Modeling Class protostellar disks in Orion

- 5,000,000 core-hour allocation to model all 200 single protostars in the VANDAM: Orion survey with NSF XSEDE supercomputers

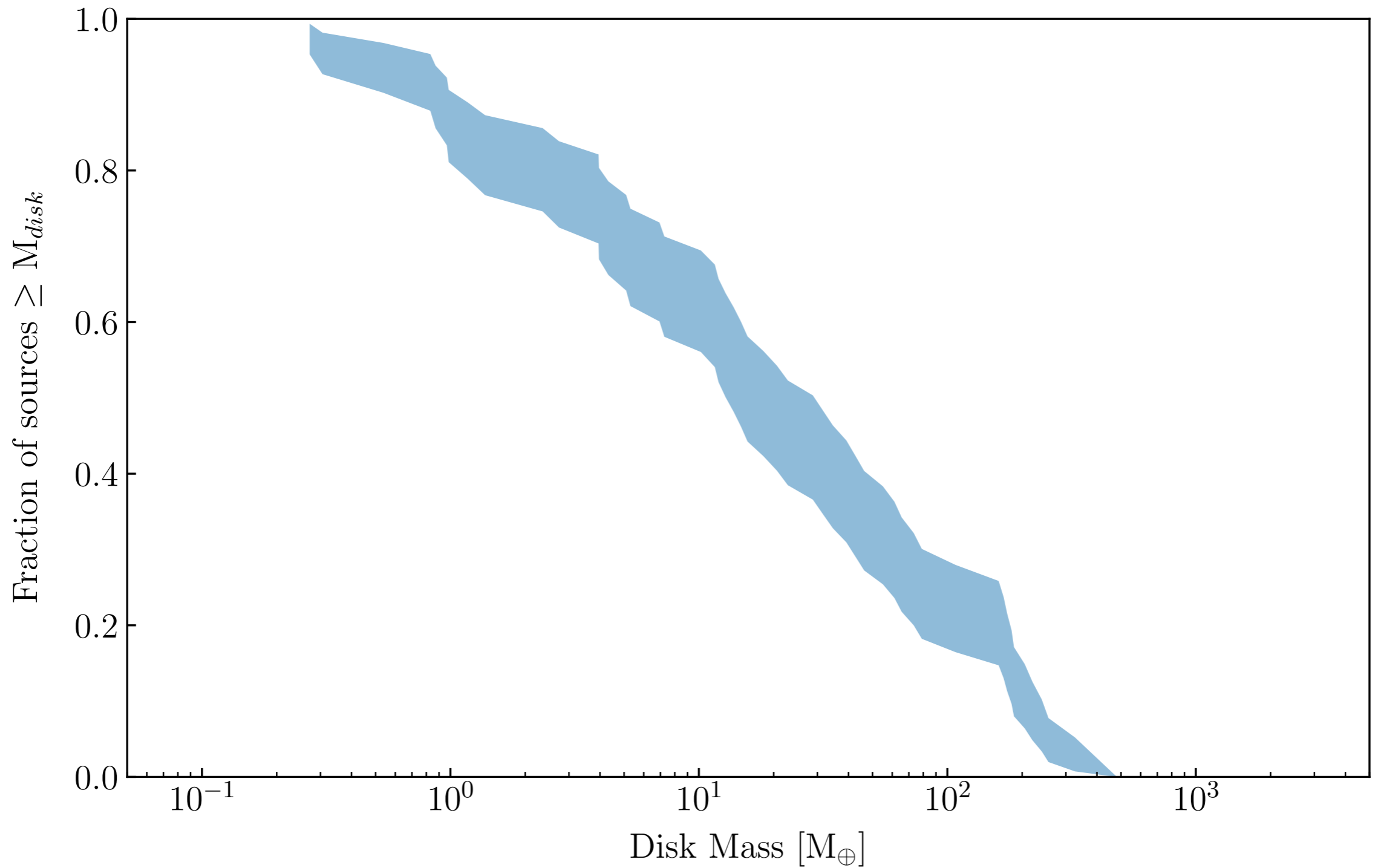


Modeling Class protostellar disks in Orion

- 5,000,000 core-hour allocation to model all 200 single protostars in the VANDAM: Orion survey with NSF XSEDE supercomputers

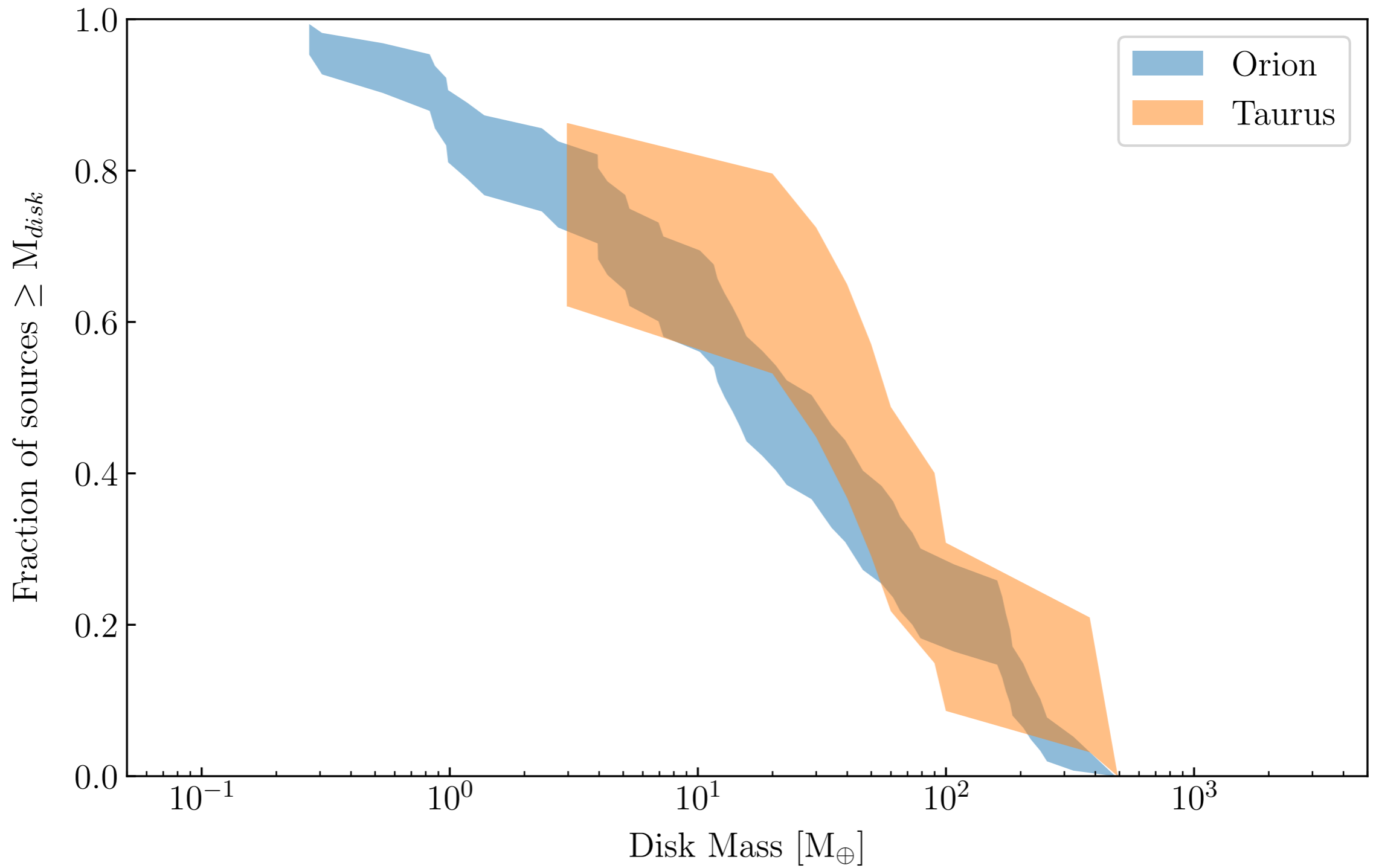


Protostellar disk demographics

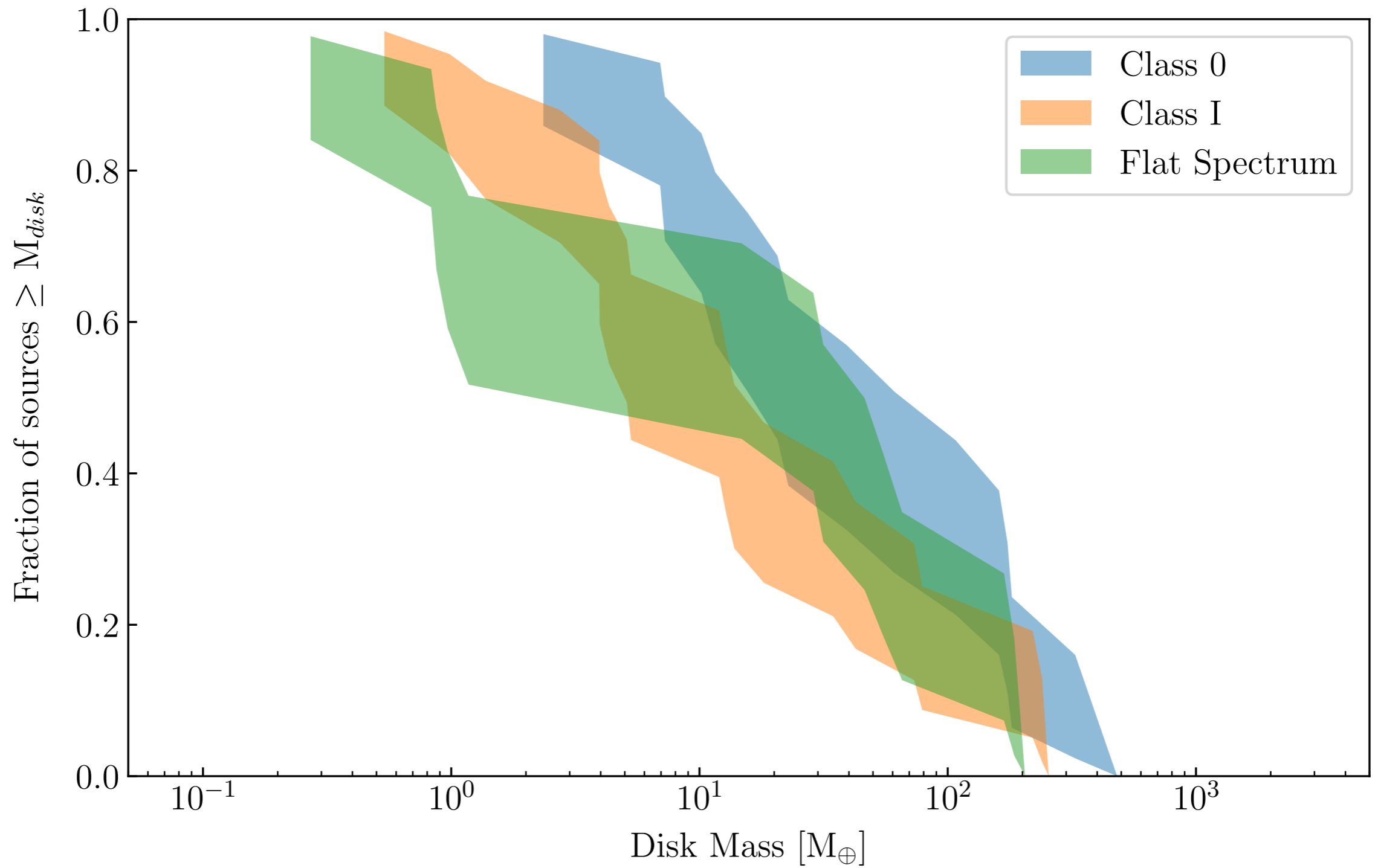


- About 50 radiative transfer models done so far

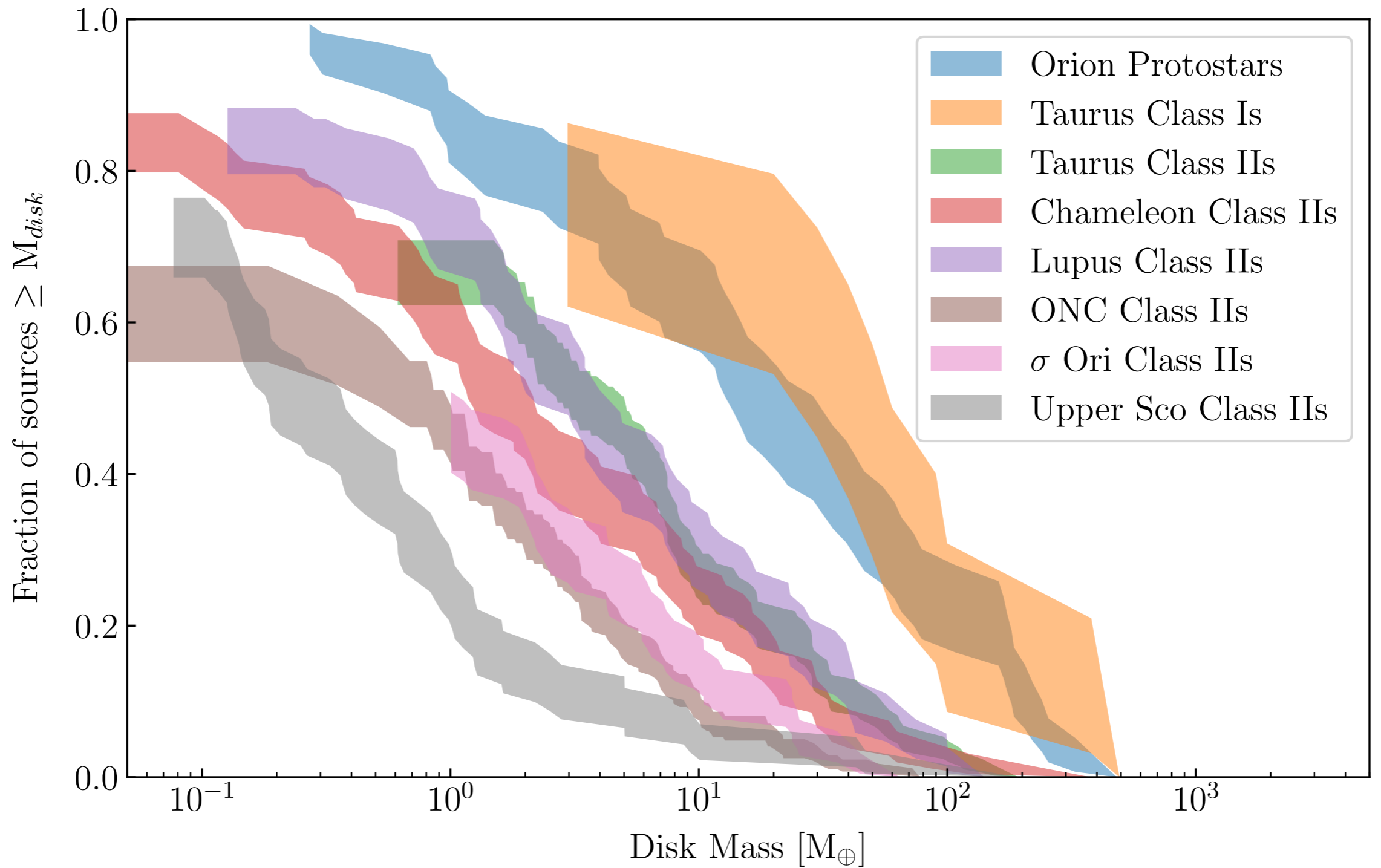
Protostellar disk demographics



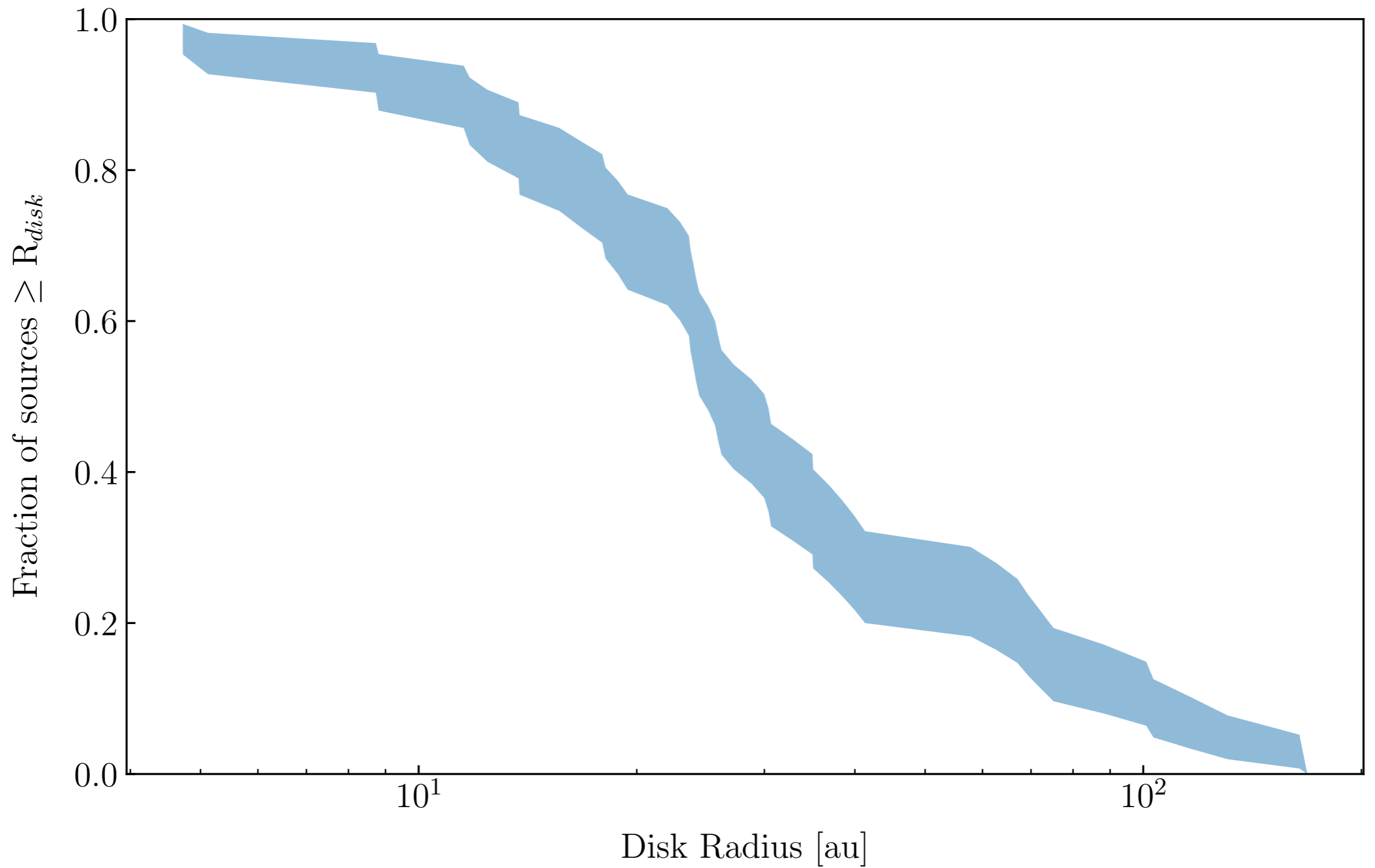
Protostellar disk demographics



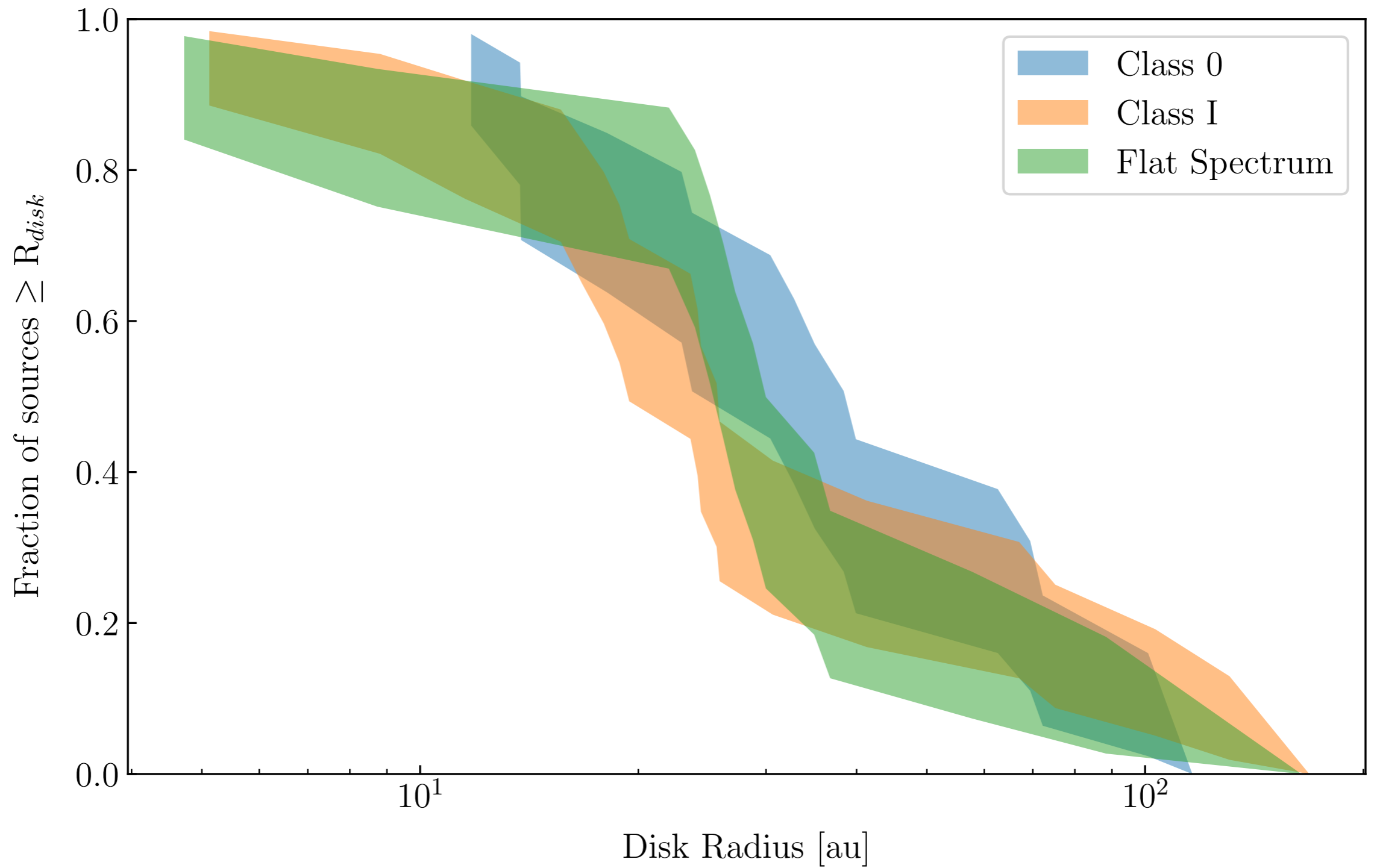
Protostellar disk demographics



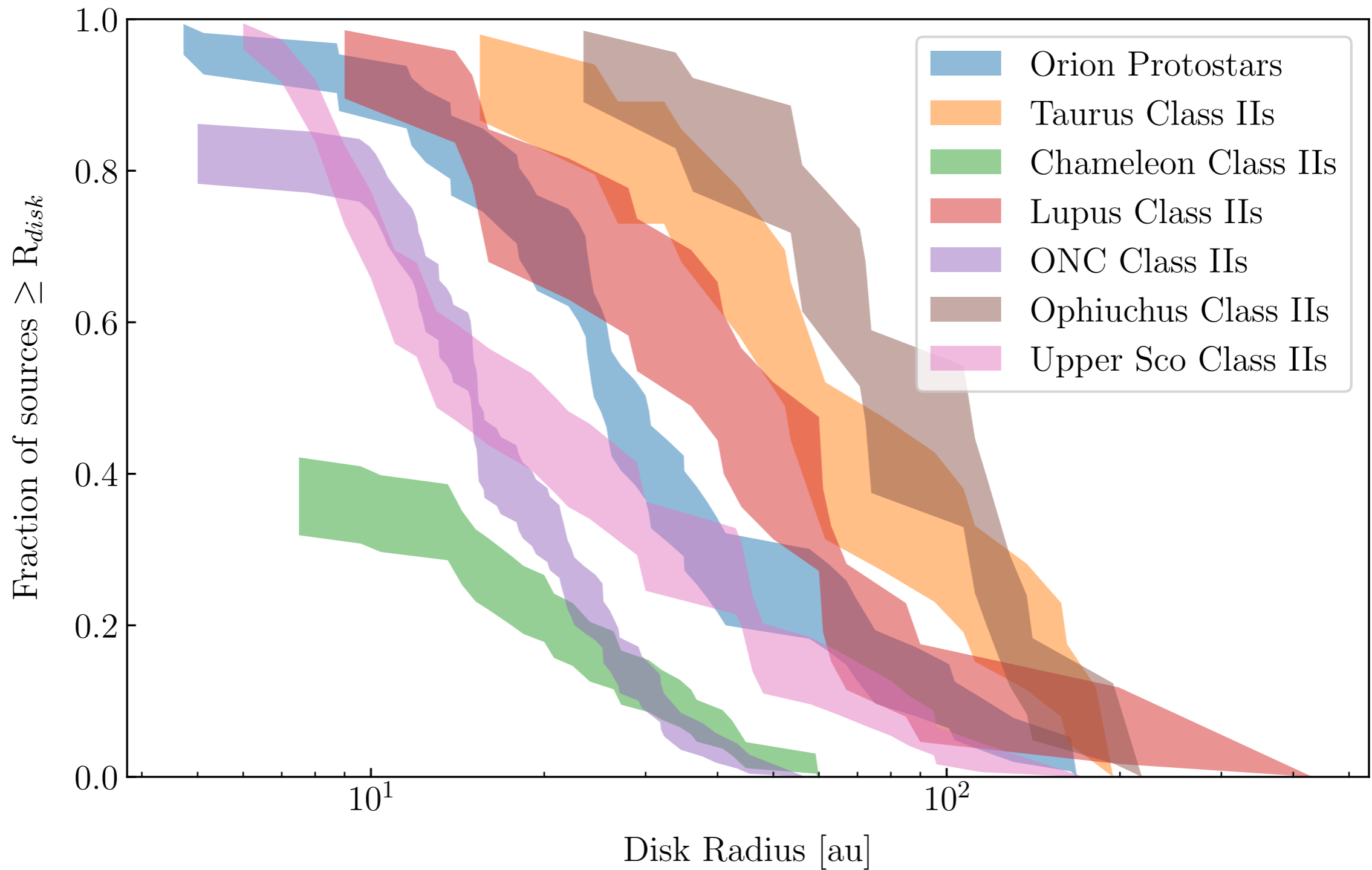
Protostellar disk demographics



Protostellar disk demographics



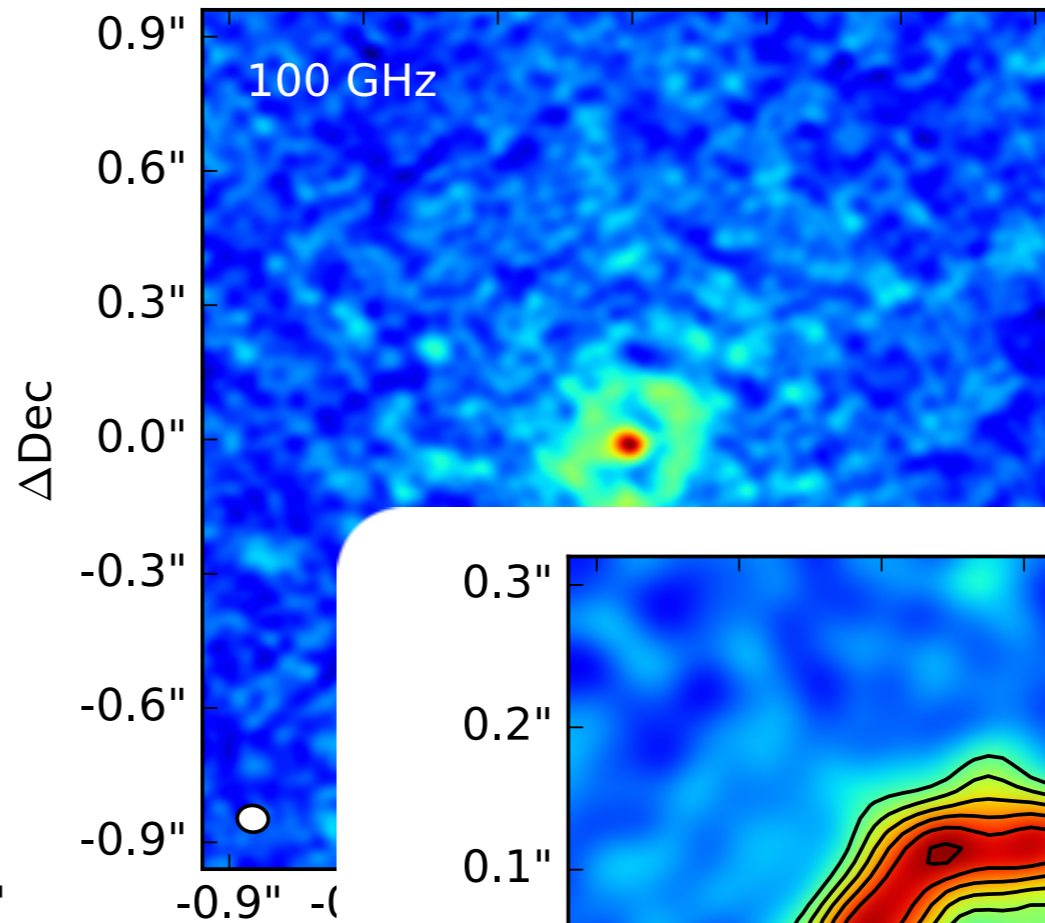
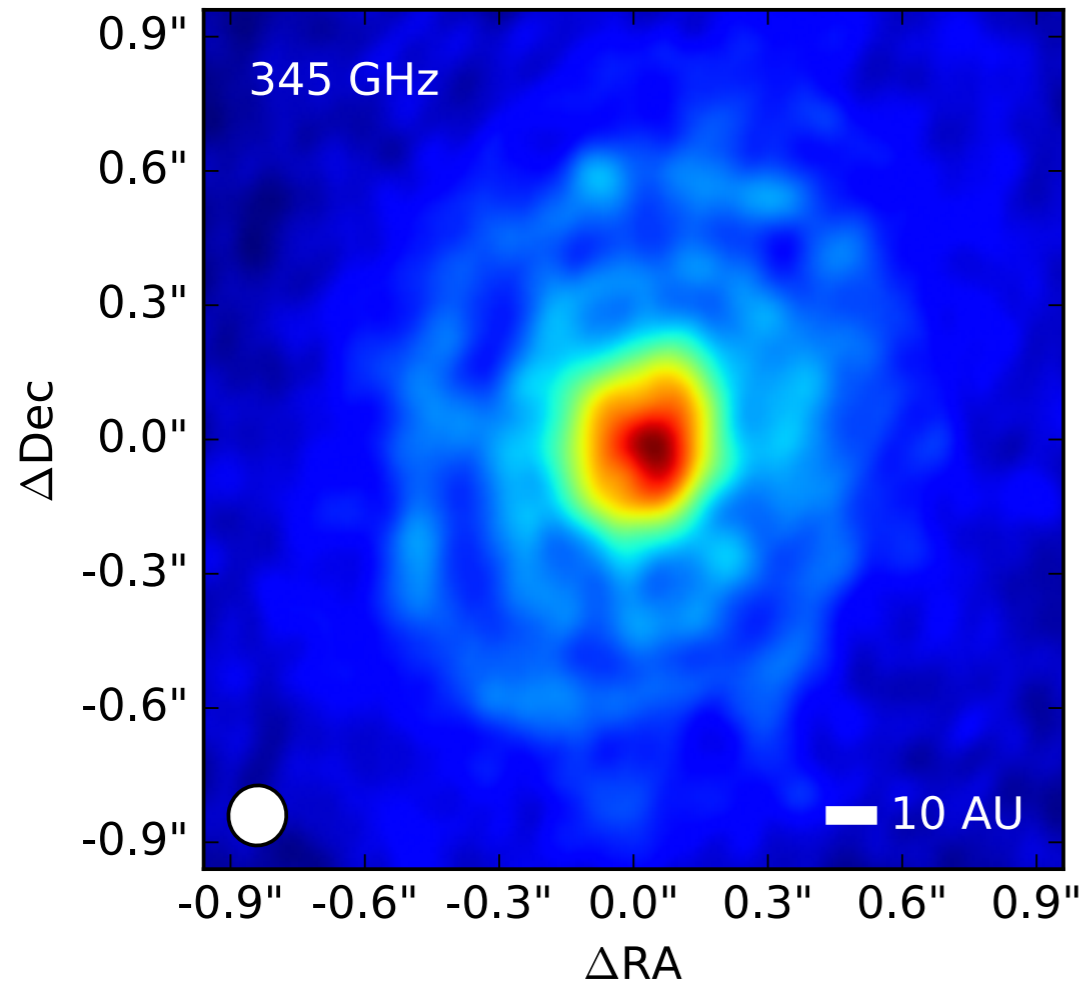
Protostellar disk demographics



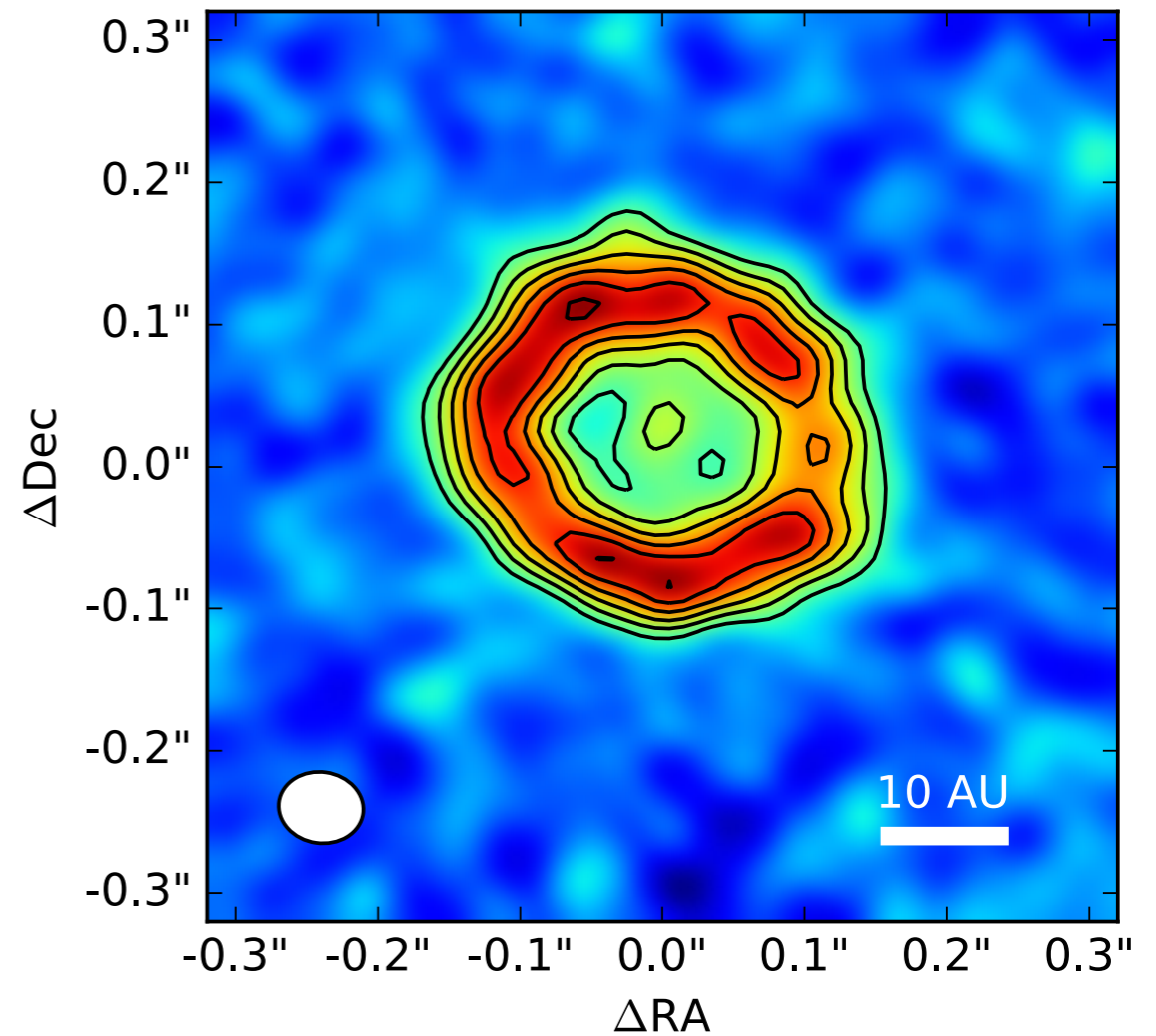
- Only 50% done => stay tuned!

Sub-structures In Embedded Disks

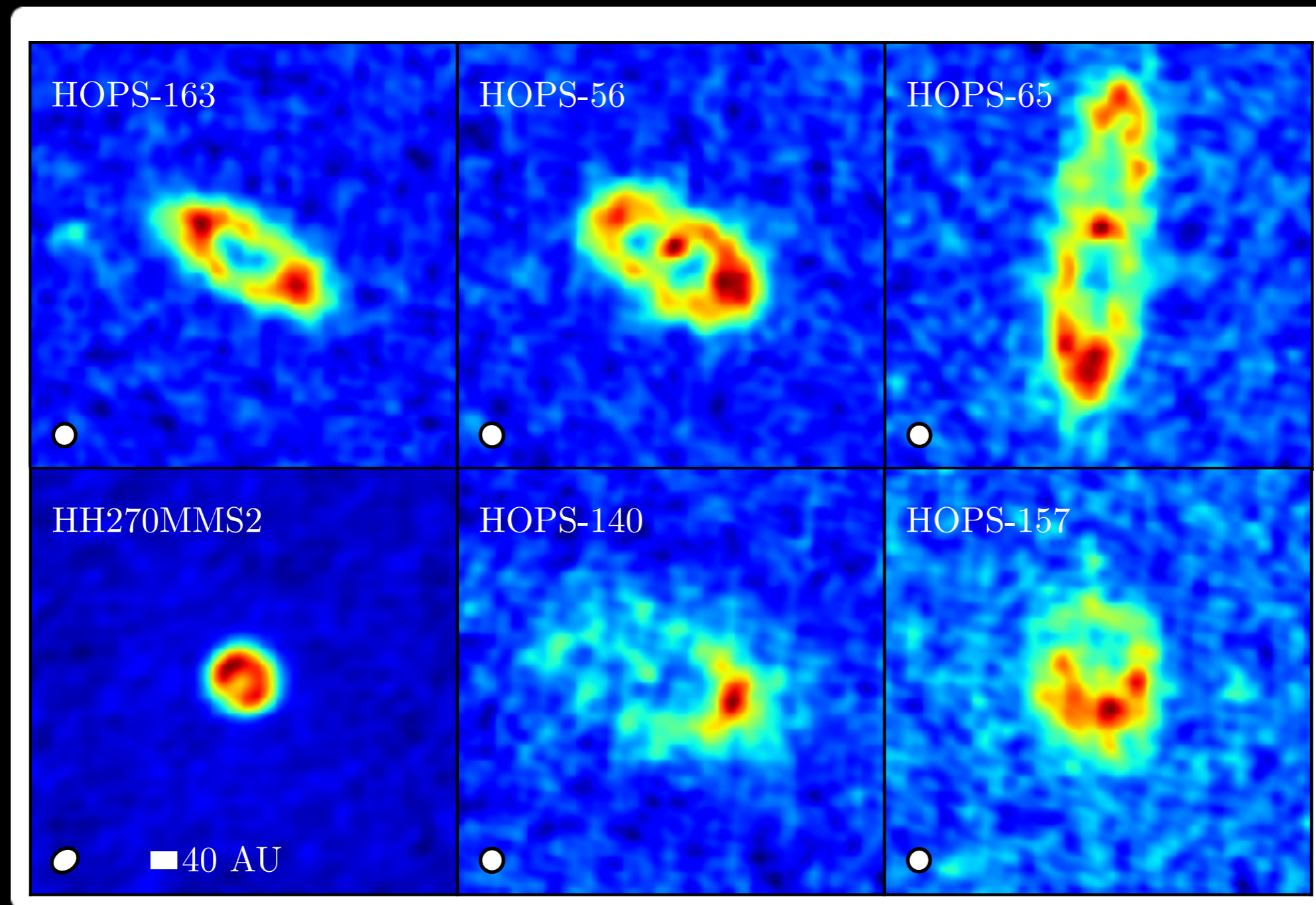
Sheehan & Eisner 2018



Sheehan & Eisner 2017a

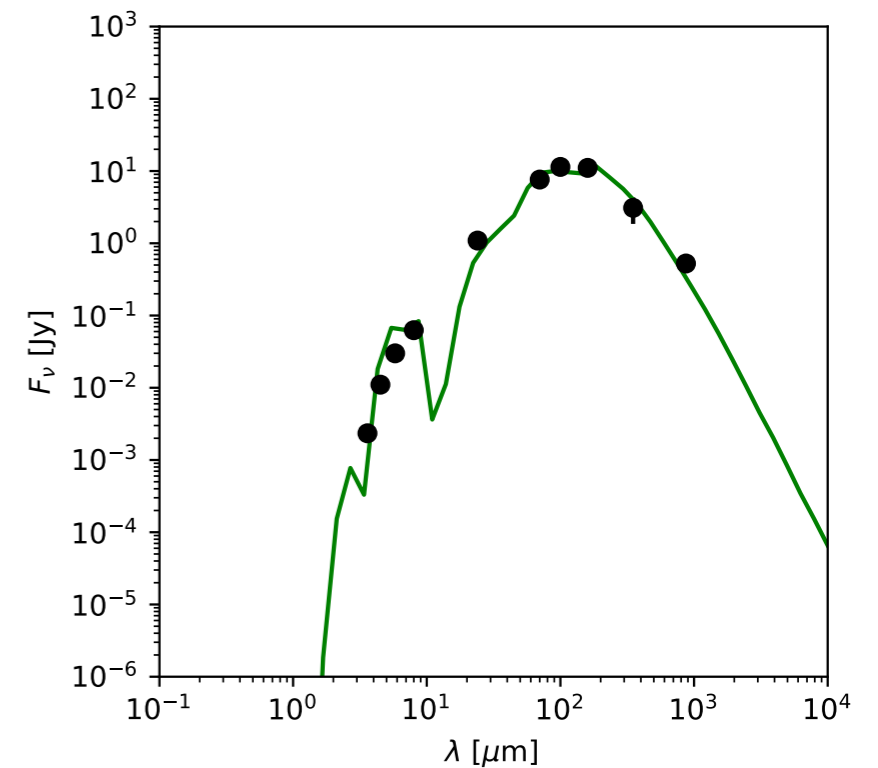
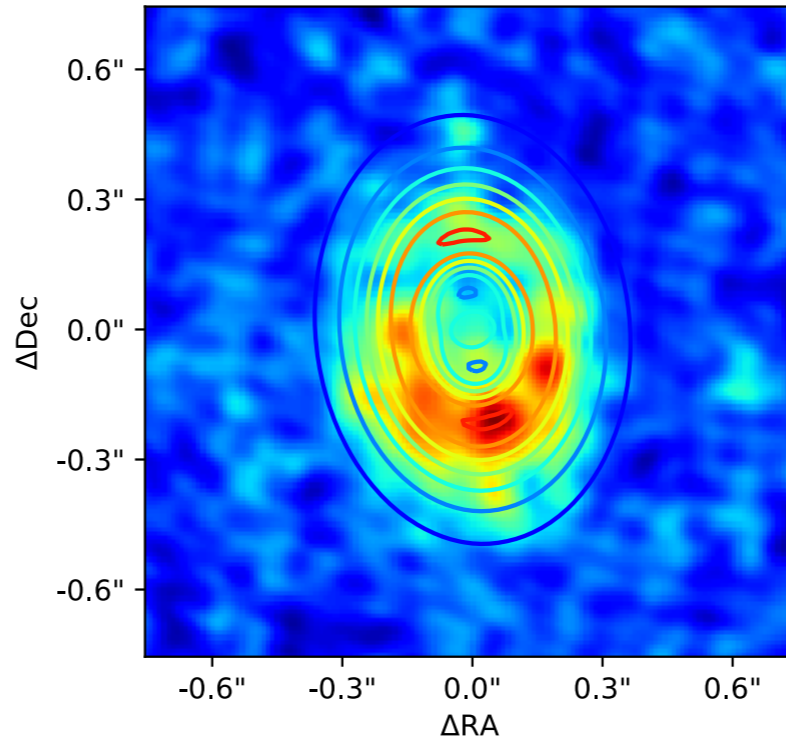
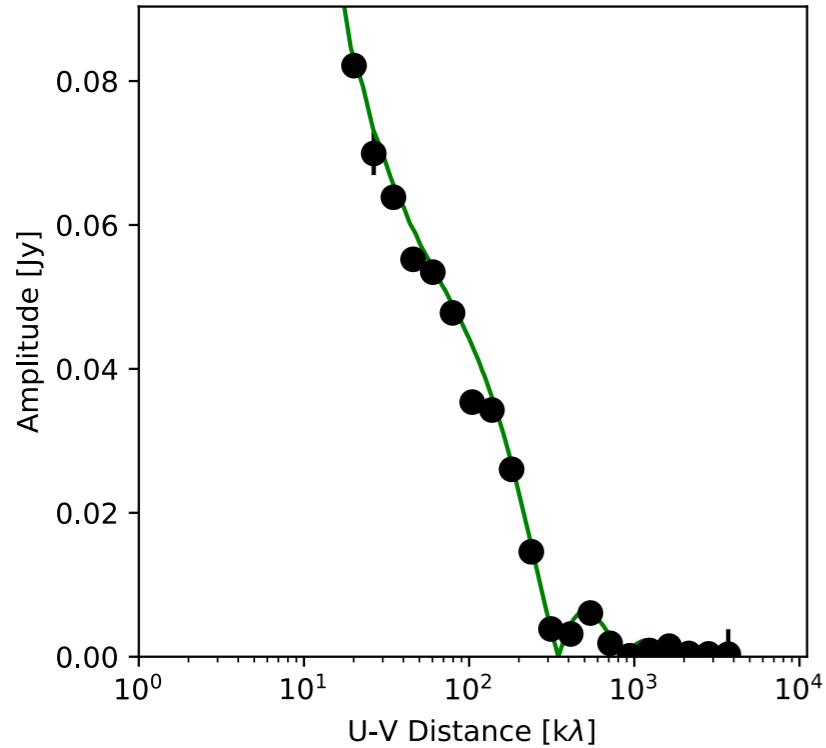
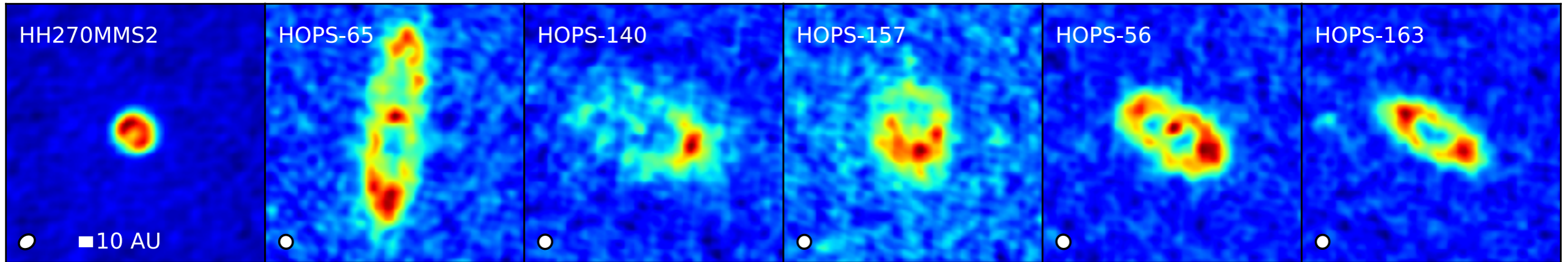


Sub-structures In Embedded Disks



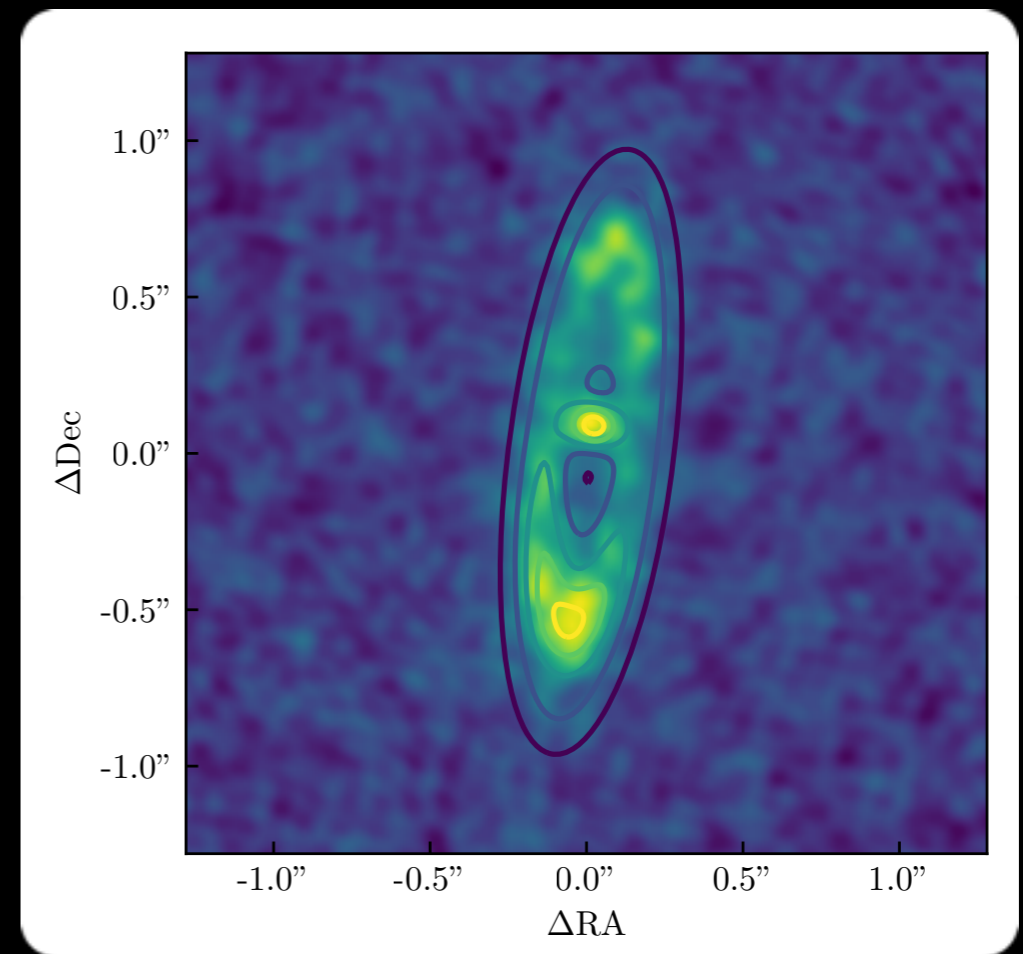
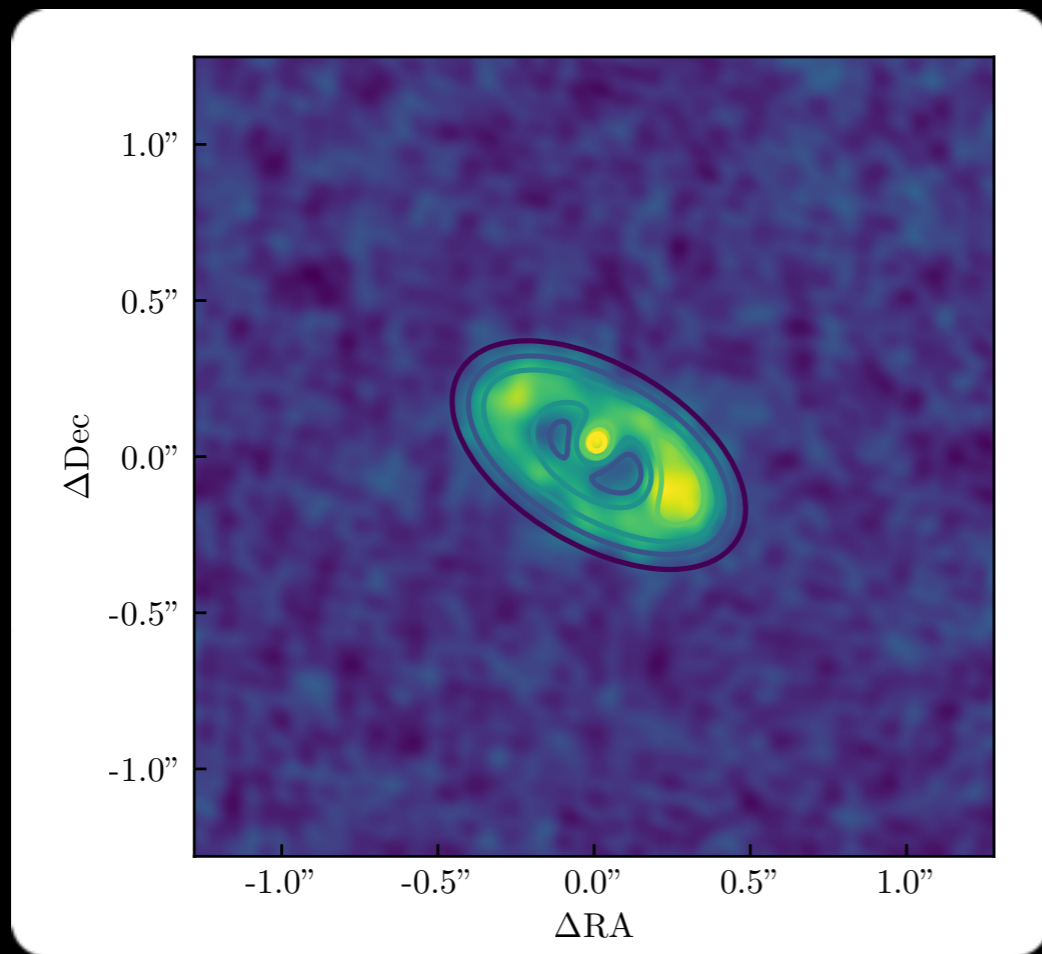
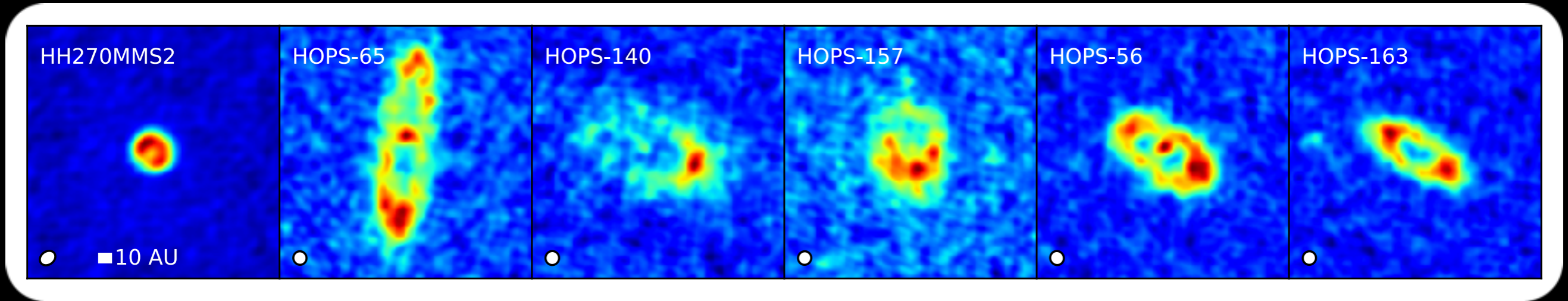
- Gap widths from 17 — 200 au
- Azimuthal asymmetries in many of the “transition disk” sources
- Several have inner disks

Sub-structures In Embedded Disks



- At least a few of the sources are heavily embedded \Rightarrow young!

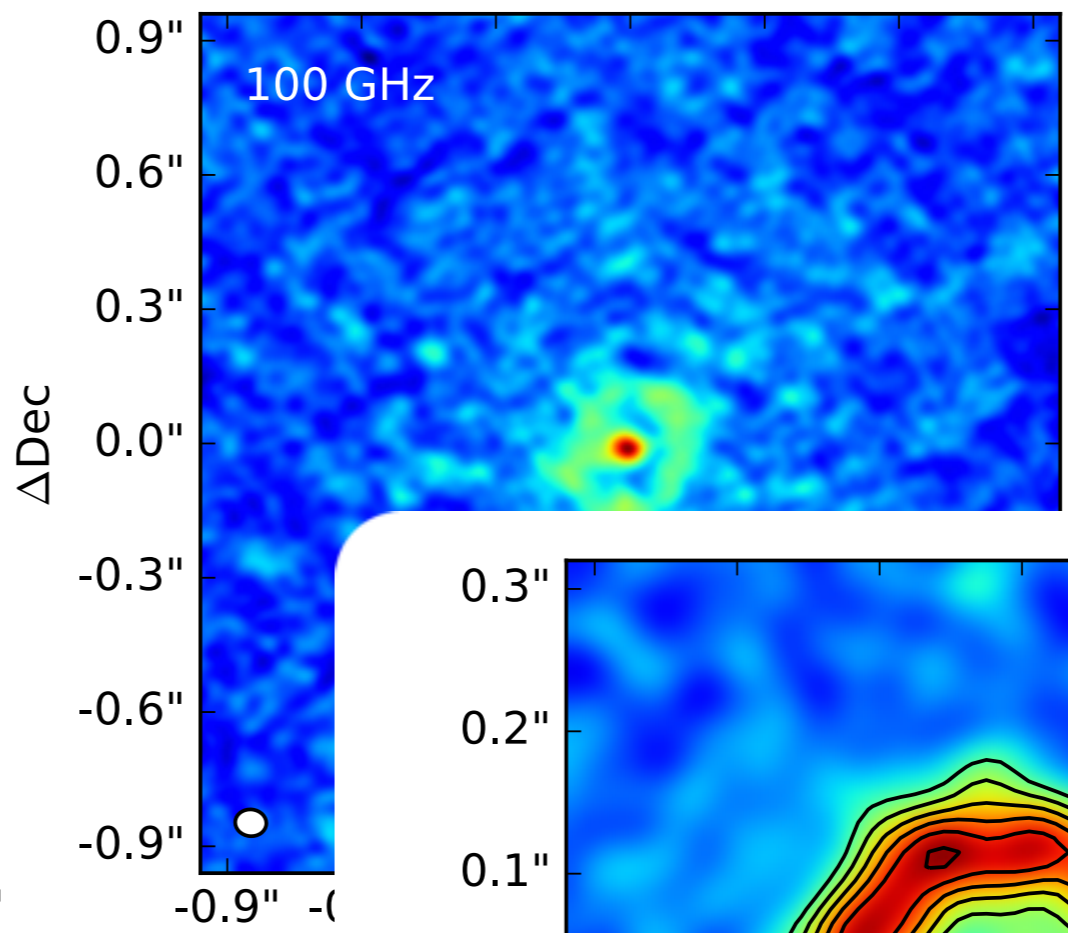
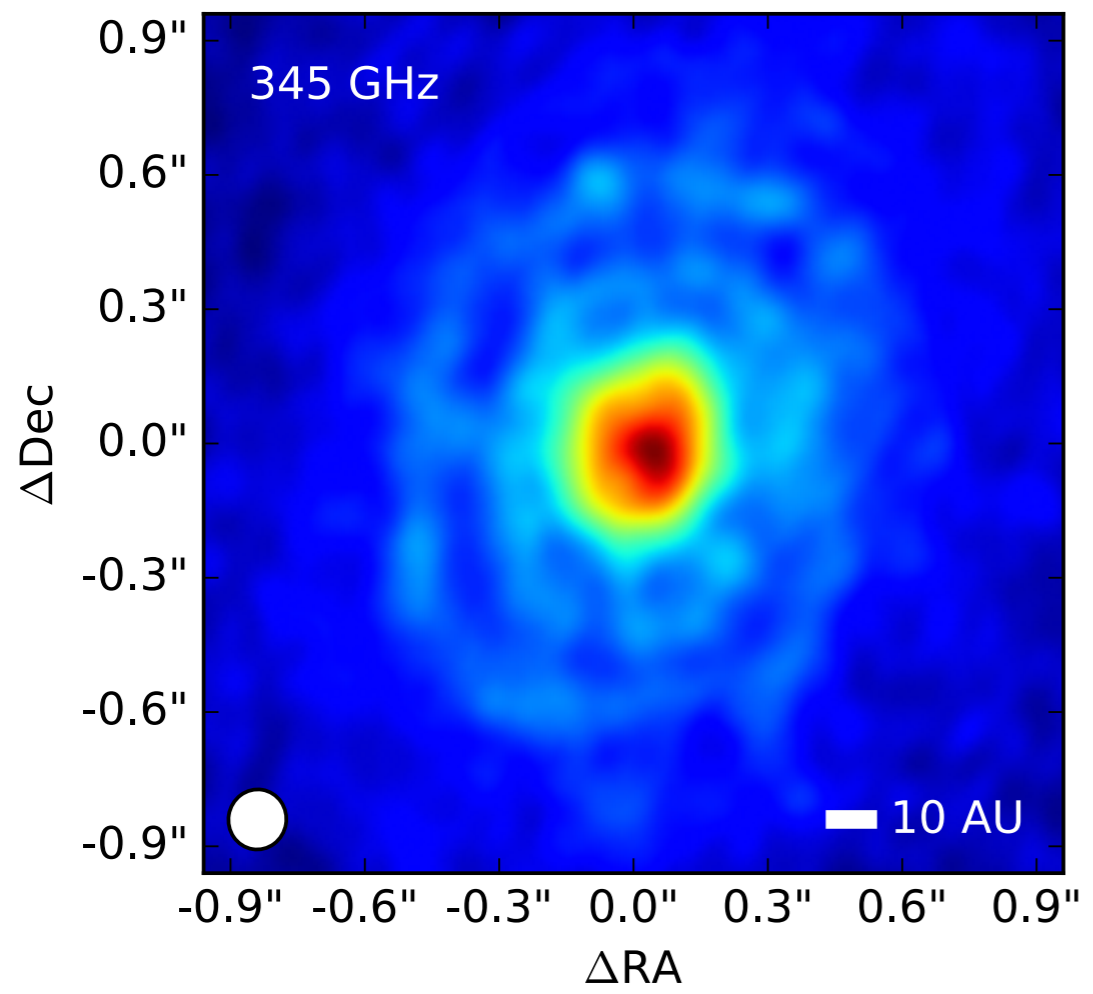
Sub-structures In Embedded Disks



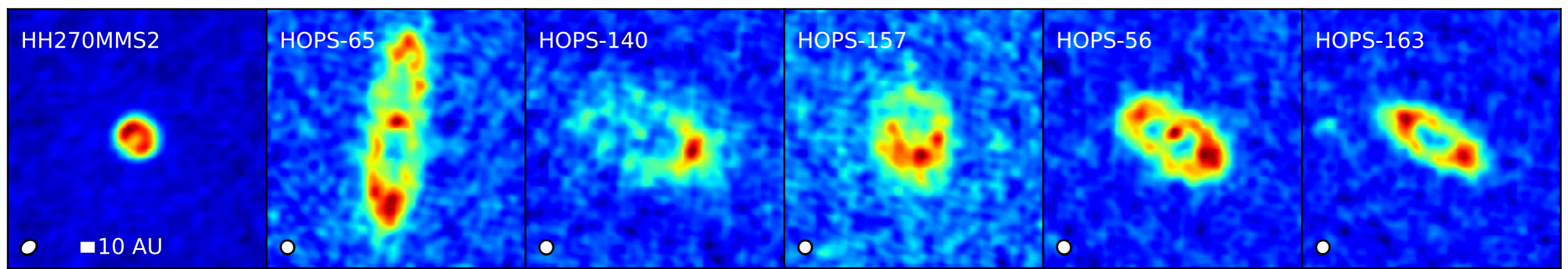
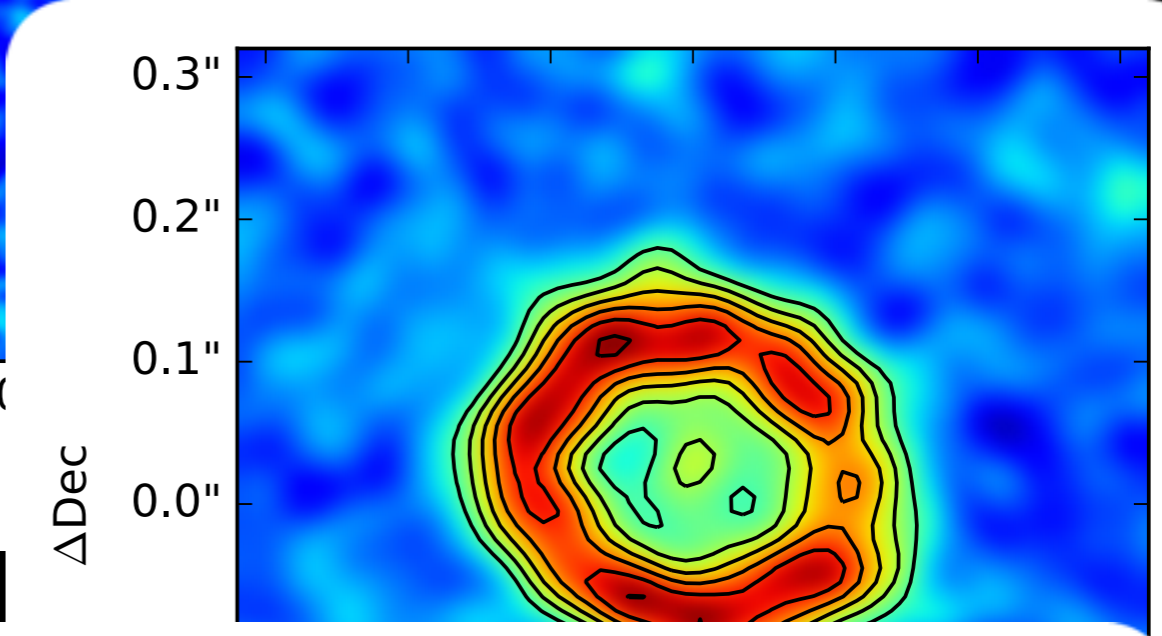
- Some of the inner disks appear to not line up with the center of the outer ring - offset by $\sim 0.05 - 0.1''$

Dynamical sculpting by large bodies?

Sheehan & Eisner 2018



Sheehan & Eisner 2017a



Questions?