

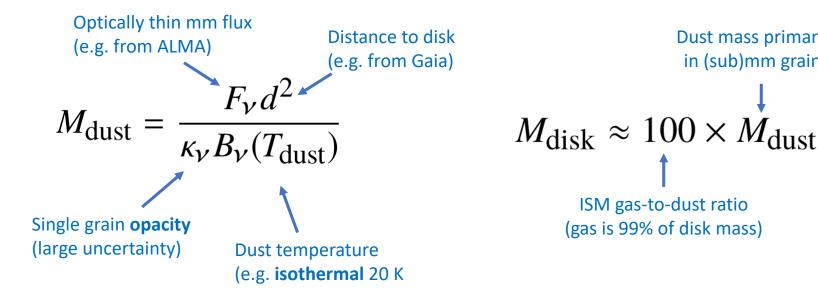
Disk masses – dust mass too low? Or an opacity issue?

Planet forming or planet hosting disks - when do we start seeing signatures of planet? Class 1? Class 0?

Populations of disks – we love spirals, gaps and rings, but what the hell is going on with all small disks? These are the norm...

Connection to exoplanets – motivation of studying disk morphologies generally connected to exoplanet demographics & architectures

Dust Disks and Total Disk Mass...

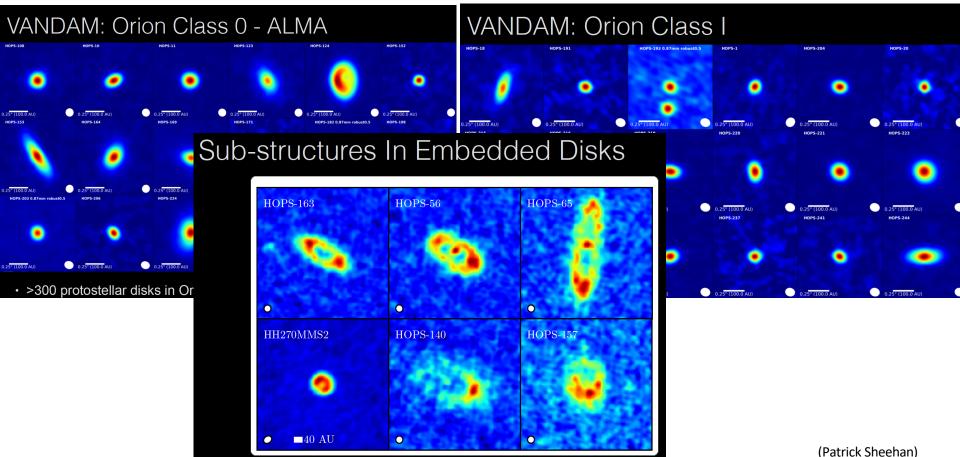


Zhaohuan Zhu – if you ignore scattering, will get the disk mass totally wrong... **Christophe Pinte** – when we use RT codes, scattering is indeed included!

Dust mass primarily

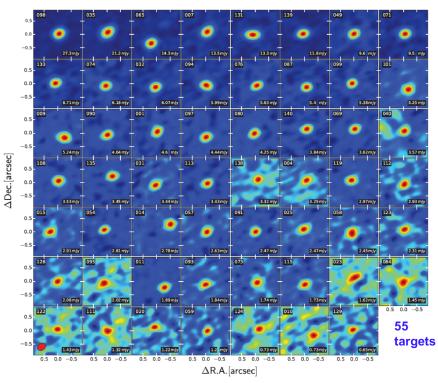
in (sub)mm grains

When do planets form?? Structure in Class I... Class 0...??



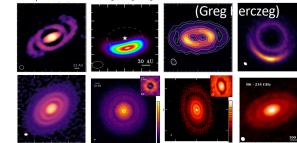
one prominent bias: targeting bright disks

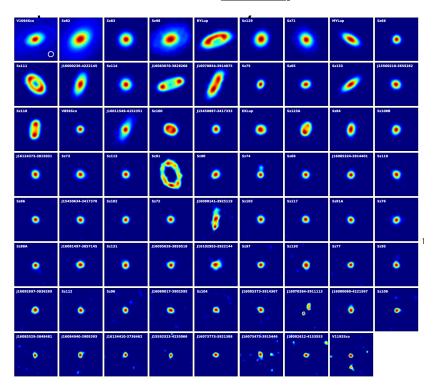
What's with all the small disks??



(Lucas Cieza: Oph ~50% unresolved disks)

(Megan Andsell: Lupus ~70% unresolved disks)





Disk structures and exoplanet populations/architectures

