# F.S.I.: FLYBY SCENE INVESTIGATION

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In memoriam Giovanni Dipierro





FONDECYT Fondo Nacional de Desarrollo Científico y Tecnológico









Blob

**S**3

0.2"≃ 28 au



### HD100453: Benisty+2017



See talks by: J. Huang L. Pérez N. van der Marel E. Ragusa M. Benisty C. Dougados A. Kóspál



# ASTROPHYSICAL CRIME SCENES (2/2)



### **RW Aur: Dai+2015**









AS 205: Kurtovic+2018

Two stars in the field

Interacting objects

Puzzling structures: bridges warped discs spirals - weird kinematics big mess

### Multi-wavelength observations are crucial

0

See talks by: L. Pérez S. Pérez R. Dong S. Facchini I. Czekala

R. Nealon



# TYPE OF SUSPECTS



## Treacherous & vicious offenders

Potentially hidden inside or outside the disc. It's likely nearby... but we do not <u>where exactly</u>



### **Observational signatures?**

Interpretation? Planet formation in multiple systems?

## Fast & strong delinquents



Should be easy to notice, but already gone... <u>Hit & run</u> ! almost never caught red handed



# ASTROPHYSICAL CLUEDO



Eccentric Dan 1.0



Inner companion eccentric and/or inclined The French Chris



Massive planet embedded in the disc Who? Massive (stellar or planetary) body
How? Weapon: gravitational torque
Modus operandi: type of orbit
Where? Inside or outside the disc
When? During the disc lifetime ~ few Myrs

### Treacherous Dan 2.0



Massive companion wandering in the outer disc regions

Fearless Valentino





Fast & furious! Extremely dangerous

# ECCENTRIC DAN 1.0: STELLAR INNER COMPANION



## Price, Cuello, Pinte et al. (2018)



Pinte, Price, Ménard et al. (2018)

# THE FRENCH CHRIS: EMBEDDED KINKY PLANET

# 1.00km/s

See talk by S. Pérez See posters by C. Pinte



# TREACHEROUS DAN 2.0: BOUND COMPANION



## Wagner, Dong, Sheehan et al. (2018) van der Plas, Ménard, Gonzalez et al. (2019)

# **DISCS & FLYBYS MODELING**

![](_page_8_Picture_2.jpeg)

q=M2/M2=1, R\_d=150 au, R\_peri=200 au inclination: β= 0° (P), 45° (P), 90° (polar), 135° (R), 180° (R)

Non-penetrating encounters:  $R_peri > R_d$ 

Degenerate problem: infinite ways to chose a flyby orbit however: orbit <-> disc structure

> **B45:** inclined prograde

**ß135:** inclined retrograde

![](_page_9_Picture_1.jpeg)

Inclined prograde (B=45°) encounter GAS – Face-on view

# FLYBYS: DYNAMICAL SIGNATURES

# FLYBYS: DYNAMICAL SIGNATURES

![](_page_10_Picture_1.jpeg)

Inclined prograde (B=45°) encounter GAS – Edge-on view

![](_page_11_Picture_0.jpeg)

10 micr

![](_page_11_Picture_2.jpeg)

## **B45** inclined prograde

1 mm •

1 micr

1 mm

l cm

t = 702 yrs •

•

### 10 micr

# l cm

## **ß135** inclined retrograde

Cuello, <u>Dipierro</u>, Mentiplay et al. (2019)

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DYNAMICS

DUST

![](_page_11_Figure_14.jpeg)

![](_page_11_Picture_15.jpeg)

# FLYBY-INDUCED ACCRETION EVENTS

![](_page_12_Figure_1.jpeg)

If we assume that  $L_* \propto M \dot{M}$ 

L suddenly x2 x10 L / with / q

-> FU Ori events?

# FLYBYS IN SCATTERED LIGHT

## Emission at 1.6 µm with MCFOST

![](_page_13_Figure_2.jpeg)

**B45** inclined prograde orbit

Cuello, <u>Mentiplay</u>, Louvet et al. (almost submittable)

# FLYBYS IN THE CONTINUUM AND CO LINES

![](_page_14_Figure_2.jpeg)

ALMA Band 7 with MCFOST and then CASA

Kinematics = information about disc orientation

Cuello, Mentiplay, <u>Louvet</u> et al. (almost submittable)

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# FLYBYS IN THROUGH THE CHANNELS

## CO channel maps for the inclined prograde case

![](_page_15_Picture_2.jpeg)

Cuello, Mentiplay, <u>Louvet</u> et al. (almost submittable)

# FLYBY VICTIM I: UX TAU

![](_page_16_Picture_2.jpeg)

Based on dynamical arguments, the orbit should be prograde. But, is it bound or unbound?

Ménard, Cuello, Gonzalez et al. (in prep)

# FLYBY VICTIM II: AS 205

1: dust continuum (Kurtovic+2018)
 2: CO mom ( (Kurtovic+2018)
 3: scattered-light obs. (S. Pérez+)
 4: preliminary hydro sim. (Guello+, in prep)
 5: CO mom 1 (Kurtovic+2018)

![](_page_17_Picture_2.jpeg)

n progre

# FLYBY VICTIMS III & IV: SR24 & FU ORI (?!)

![](_page_18_Figure_1.jpeg)

![](_page_18_Picture_2.jpeg)

### S. Pérez, Hales, Liu et al. (submitted)

In SR24: bridge between both stars surprisingly no mm disc around SR24N

In FU Ori: non-coplanar spiral excellent match with scattered light spiral! –> non-coplanar interaction!

![](_page_18_Picture_7.jpeg)

# WHEN SHOULD YOU SUSPECT A FLYBY?

<b>PROGRADE?</b>	<b>OBSERVATIONAL SIGNATURES</b>	<b>RETROGRADE?</b>
Yes	Bridges & spirals with large pitch angles	~
Yes	Dust trapping in the spirals	Yes
Yes	<b>Disc truncation</b>	No
~	Warped discs	Yes
Yes	Diffuse halo	No
Yes	Accretion events / outbursts ?	No

Cuello, Mentiplay, Louvet et al. (almost submittable)

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# FINAL WORDS

## **Disc structure <-> Flyby parameters**

- 1)
- 2)
- 3)
- 4)
- 5)

![](_page_20_Picture_7.jpeg)

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Do we observe stellar flybys? How does the disc react to a flyby? How does a flyby look like? Can we identify flyby suspects? Implications for planet formation?

### **OPEN QUESTIONS**

disc long-term evolution after the encounter? planet formation in perturbed/multiple systems? - connexion with the Solar & exoplanetary systems?

14 **DRINKS** 

![](_page_20_Picture_17.jpeg)